

Workshop on Precision Measurements at Low Energy

Neutron Decay and Neutron Oscillations

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A short history of unification

'Unification' in physics:

Unified description for previously disparate phenomena

century:

unification of:

17th-18th "Himmelsmechanik" and earthly mechanics → MECHANICS

19th Magnetism, electricity, optics, heat radiation → ELECTRODYNAMICS

Mechanics and thermodynamics → STATISTICAL MECHANICS

20th/1: Quantum mechanics and electrodynamics → QED

/2: QED, weak interact., QCD → STANDARD MODEL of Particle-Physics

21st/1: Particle physics and cosmology,

Quantum mechanics and relativity: → ???

Standard Model of particle physics ...

small input:

Symmetries
(‘principia’)



Gauge principle: $\psi'(x) = e^{i\varphi(x)}\psi(x)$
applied to $U(1)\times SU(2)\times SU(3)$,

Lorentz invariance: $x' = L\cdot x$,

CPT, T, ... invariance, ...

rich output:

Interactions:

→ equations of motion

Maxwell,
Schrödinger,
Dirac, ...

basis for:
technology,
chemistry,
molec. biology, ...

→ existence of photons, gluons, W^\pm , Z^0 (carriers of interaction)

→ conservation of charges (sources of interaction)

...is very successful ...

... but is not complete:

Unsolved problems:

3 particle families

9 masses (+3 m_ν ?)

4 quark-phases (+4 lepton-phases?) ...

gravitation/relativity

baryon-asymmetry of universe

→ mass-energy content of universe ...

Is there one 'unified' solution to all remaining problems ?

If so, this unification will probably take place

at extremely high energies.

"Precision measurements at low energy"

Why search then at low energy?

When an unknown process exists

at an unreachable energy scale $M \sim 10^{5\dots 19}$ GeV,

then the propagator is **the same**,

whether one works at 1 neV or at 10^3 GeV, and is small:

e.g. $1/(p^2+M^2) \rightarrow 1/M^2$

needed is: **HIGH PRECISION,**

which is highest **at LOW ENERGY**

Low energy: (mainly) 1st family

	quarks		leptons	
3 rd :	b	t	τ	ν_τ
2 nd :	s	c	μ	ν_μ
1 st :	d	u	e	ν_e

first family is:

- abundant,
- long-lived,
- useful.

Precision tests of Standard Model

To test a theory needs both:

- precise theoretical predictions
- precise experimental results

Both is feasible for **electroweak** and for **gravitational** interaction


Neutron **gravitational** interaction:

- limits on $m_{\text{grav}} - m_{\text{inert}}$
- in neutron self-interference
- neutron quantization in gravitational field \rightarrow limits on extra dimensions

Neutron **electroweak** interaction:

- electric properties: limits on neutron charge, neutron EDM, ...
- neutron-nuclear weak interaction: \rightarrow new information on strong interaction
- **neutron decay: this talk**

Precision reached in neutron work

- in energy: $\delta E = \pm 10^{-23}$ eV = $\pm 0.000\ 000\ 000\ 000\ 000\ 000\ 000\ 01$ eV
reached in neutron oscillations and neutron EDM
- in momentum: $\delta p/p = \pm 10^{-11}$: 1 Å/10m
reached in neutron charge exp't. 
- in polarization: $\delta P = \pm 10^{-7}$
reached in neutron-nuclear PNC interaction

1. Neutron β -decay

	quarks		leptons		
3 rd :	b	t	τ	ν_τ	} high energy
2 nd :	s	c	μ	ν_μ	
1 st :	d \rightarrow u		+ e + ν'_e		neutron decay

$$n \rightarrow p^+ + e^- + \nu'_e$$

neutron lifetime $\tau \approx 15$ min

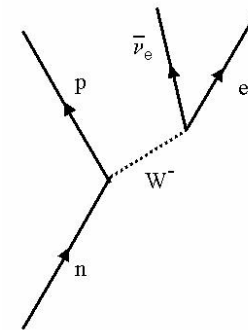
β -endpoint energy: $E_{\max} = 782$ keV

V–A Standard Model:

two coupling constants ($E \rightarrow 0$):

vector c.c. g_V

axial-vector c.c. g_A



Neutron decay data are useful

Many processes have the same Feynman diagram as neutron decay:

Primordial element formation
(^2H , ^3He , ^4He , ^7Li , ...)

$$n + e^+ \rightarrow p + \nu'_e \quad \sigma_\nu \sim 1/\tau$$

$$p + e^- \rightarrow n + \nu_e \quad \sigma_\nu \sim 1/\tau$$

$$n \rightarrow p + e^- + \nu'_e \quad \tau$$

Solar cycle

$$p + p \rightarrow ^2\text{H} + e^+ + \nu_e$$

$$p + p + e^- \rightarrow ^2\text{H} + \nu_e \text{ etc.} \sim (g_A/g_V)^5$$

Neutron star formation

$$p + e^- \rightarrow n + \nu_e$$

Pion decay

$$\pi^- \rightarrow \pi^0 + e^- + \nu'_e$$

Neutrino detectors

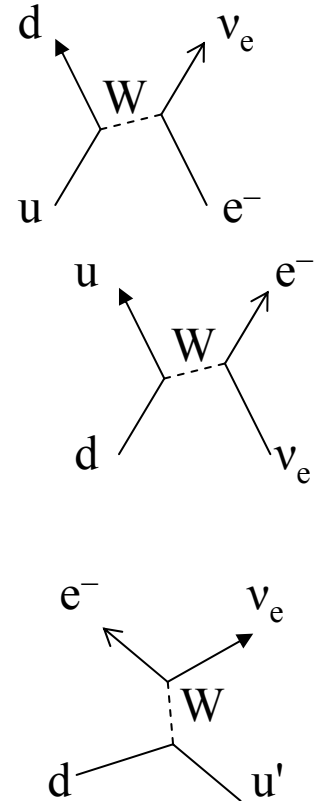
$$\nu'_e + p \rightarrow e^+ + n$$

Neutrino forward scattering

$$\nu_e + n \rightarrow e^- + p \text{ etc.}$$

W and Z production

$$u' + d \rightarrow W^- \rightarrow e^- + \nu'_e \text{ etc.}$$



... precision data of weak interaction parameters

today only from neutron decay ('moving target')

Only few n-decay parameters in Standard Model ...

$$\begin{aligned} \text{n-decay rate: } \quad \tau^{-1} &= \text{const} (|g_V|^2 + 3|g_A|^2) \\ &= \text{const} G_F^2 |V_{ud}|^2 (1+3|\lambda|^2) \end{aligned}$$

with: $|g_V|^2 = G_F^2 |V_{ud}|^2$ Fermi c.c. G_F from μ -decay.

At most 3 parameters needed: V_{ud} , λ , φ :

- CKM matrix element V_{ud} ,
- ratio of c.c. $\lambda = g_A/g_V = |\lambda|e^{i\varphi}$
- relative phase φ
(T-symmetry: $\varphi = 180^\circ$)

... but many n-decay observables:

PDG 2006:

decay rate $\tau^{-1} = \text{const} |V_{ud}|^2 (1 + 3\lambda^2)$, $\tau = 885.7(8)$ s (talk S. Paul) (re-)measured by:

e - $\bar{\nu}_e$ - correl.: $a = \frac{1 - \lambda^2}{1 + 3\lambda^2} = -0.103(4)$ (talk N. Severijns) *a*SPECT

β - asymmetry: $A = -2 \frac{\lambda(\lambda + 1)}{1 + 3\lambda^2} = -0.1173(13)$ sensitive to λ , rhc phase ζ

$\bar{\nu}_e$ - asymmetry: $B = 2 \frac{\lambda(\lambda - 1)}{1 + 3\lambda^2} = 0.981(4)$ sensitive to mass of W_R } PERKEO

p - asymmetry: $C = 0.2748 \frac{4\lambda}{1 + 3\lambda^2} = -0.238(11)$

triple correlations: $D = -2 \frac{\lambda \sin \varphi}{1 + 3\lambda^2} = -4(6) \cdot 10^{-4}$ T - viol. TRINE

$R = \dots$, etc. T - viol. (talk N. Severijns)

problem is overdetermined:

many tests of Standard Model possible

Therefore many derived quantities ...

Standard model parameters:

axial to vector coupling c.c.

$$\lambda = g_A/g_V \text{ from } \beta\text{-asymmetry } A \text{ (or } a)$$

CKM- matrix element $|V_{ud}|$ from:

$$|V_{ud}|^2 = \frac{(4908.7 \pm 1.9) \text{ sec}}{\tau \cdot (1 + 3\lambda^2)}$$

unitarity test of CKM-matrix

$$\Delta = V_{ud}^2 + V_{us}^2 + V_{ub}^2 - 1 = 0 ?$$

phase between g_A and g_V :

$$\varphi = 180.06(7)^\circ$$

i.e.

$$e^{i\varphi} = -1 \text{ (T-symmetric V-A)}$$

weak magnetism, ...

further derived quantities:

all ν -p, ... weak cross-sections

$$\sigma_{\nu p}/E_\nu = 0.67 \cdot 10^{-38} \text{ cm}^2/\text{GeV}$$

number of ν -families

$$N_\nu = 2.5 \pm 0.6$$

baryonic matter in universe

$$\rho/\rho_{\text{crit}} = (3.3 \pm 0.7) \%$$

... more derived quantities

beyond Standard model:

mass of right-handed boson $m(W_R) > 300 \text{ GeV}/c^2$ (90% c.l.)

left-right mixing angle $-0.20 < \zeta < 0.07$ (90% c.l.)

scalar weak interaction amplitudes g_S

tensor weak interaction amplitudes g_T

Fiertz interference amplitude b

second class amplitudes

neutrino helicity < 1 ? (semileptonic decays)

... and others

Aim: measure all these parameters to the highest precision possible !

CKM quark mixing

μ -decay: $\sim G_F^2$

n-decay: $\sim 0.95 \cdot G_F^2 = \cos^2 \theta_C \cdot G_F^2$ ($\Delta S=0$), strangeness S

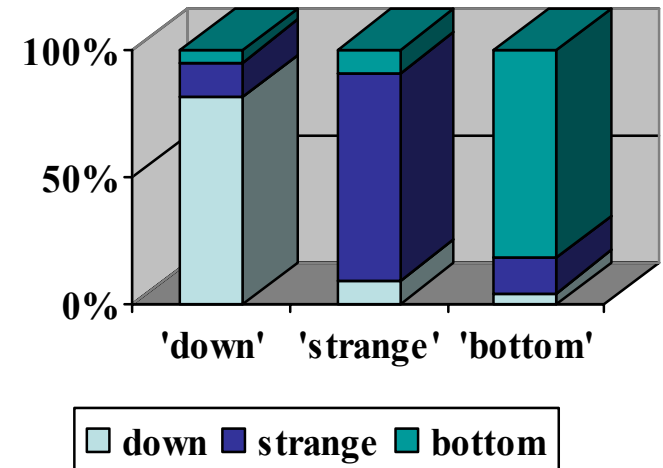
K-decay: $\sim 0.05 \cdot G_F^2 = \sin^2 \theta_C \cdot G_F^2$ ($\Delta S=1$), Cabbibo angle θ_C

three generation Standard Model:
under the weak interaction,
quarks are not the usual quarks, but:

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

↑ weak eigenstates

↑ mass eigenstates



quark mixing is **pure rotation in flavor space** ('zero-sum game'),
Cabbibo-Kabojashi-Maskawa quark mixing matrix is **unitary**

$|V_{ud}|$ from neutron decay

from $\tau^{-1} = \text{const } G_F^2 |V_{ud}|^2 (1+3|\lambda|^2)$:

$$|V_{ud}|^2 = \frac{(4908 \pm 4) \text{sec}}{\tau \cdot (1 + 3\lambda^2)}$$

i.e. two measurements needed: τ and λ

History of neutron lifetime

PDG:

- neutrons 'in-beam': 1960: $\tau = (1010 \pm 30) \text{ s}$

1982: $\tau = (925 \pm 11) \text{ s}$

- stored UCN: 1989: $\tau = (888 \pm 3) \text{ s}$

PDG since 2002: $\tau = (885.7 \pm 0.8) \text{ s}$

2005: $\tau = (878.5 \pm 0.8) \text{ s} ?$

History of $\lambda = g_A/g_V$

PDG: from β -asymmetry A in neutron-decay:

1960: $\lambda = -1.19(2)$

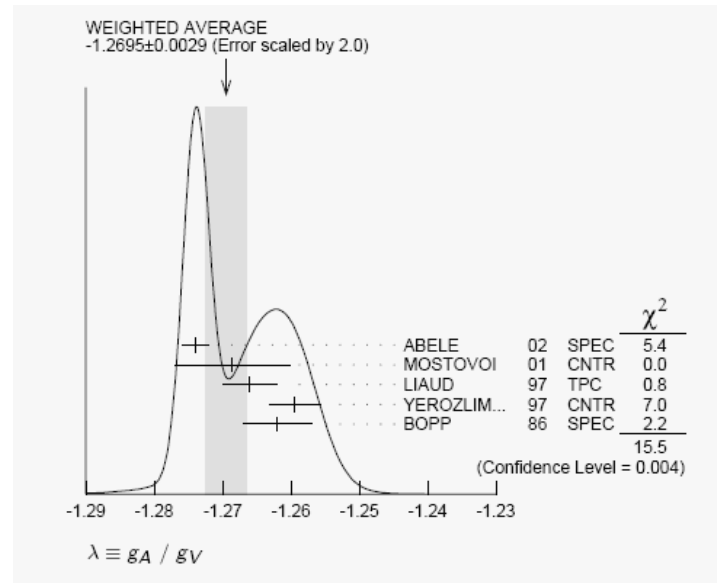
1975: $\lambda = -1.25(2)$

1994: $\lambda = -1.261(4)$

PDG since 2002: $\lambda = -1.270(3)$

sys. corrections: $>15\%$

sys. corrections: $5 \cdot 10^{-3}$



Is the unitarity of the CKM matrix violated in neutron β -decay?

H. Abele et al., PRL 88, 211801 (2002)

CKM-Unitarity of first row:

$$\begin{aligned} |V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 &= 1 - \Delta \\ &\quad \uparrow 0.0000 \\ &\approx \cos^2\theta_C + \sin^2\theta_C \end{aligned}$$

Standard Model: unitarity, i.e. $\Delta = 0$

Using the world average of neutron lifetime: $\tau = 885.7(7)$ s,

almost all error came from $\lambda = g_A/g_V$.

λ from new measurement of β -asymmetry: $A = -\frac{2\lambda(\lambda+1)}{1+3\lambda^2} \cong -0.1189(7)$

Deviation from unitarity (Perkeo)

from: $A = -0.1189(7)$

→ $\lambda = g_A/g_V = -1.2739(9)$

→ $|V_{ud}| = 0.9713(13)$ neutron

in detail: $= 0.9717(4)_\tau (12)_\lambda (4)_{\text{rad.corr.}}$

↑ culprit?

with : $|V_{us}| = 0.2196(23)$ Kaon

$|V_{ub}| = 0.0036(9)$ B-meson (PDG 2002)

$$\Delta = |V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 - 1:$$

$$\Delta = 0.0084 \pm 0.0028$$

3σ deviation from zero

(PERKEO 2002)

Similar results from other β -decays:

H. Abele, E. Barberio, D. Dubbers, F. Glück, J.C. Hardy,
W.J. Marciano, A. Serebrov, N. Severijns, Eur. Phys. J. C33, 1 (2004)

$$|V_{ud}| = 0.9717 \pm 0.0013 \quad \text{neutron}$$

$$|V_{ud}| = 0.9738 \pm 0.0004 \quad 0^+ \rightarrow 0^+ \text{ nuclei}$$

$$|V_{ud}| = 0.9728 \pm 0.0030 \quad \text{pion}$$

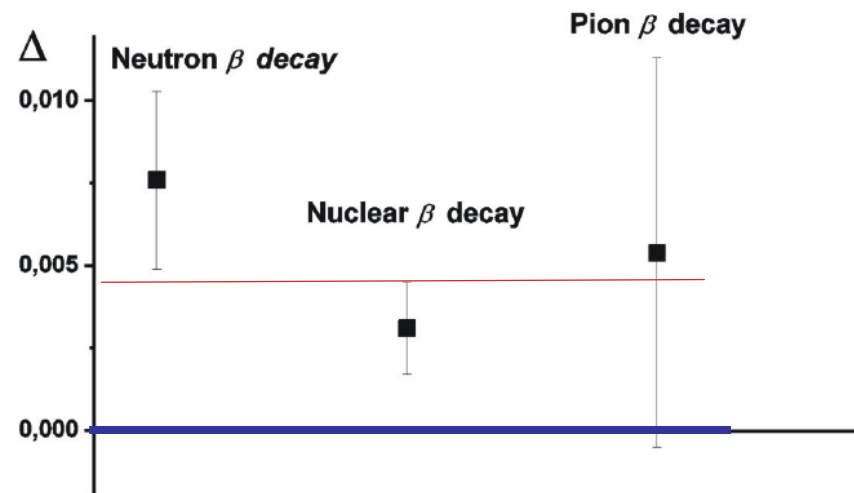
PDG 2004:

combined: $\Delta = 0.0040 \pm 0.0012$ (all V_{ud})

if $\Delta \neq 0$ is due to right-handed currents:

phase $\zeta = 0.0020 \pm 0.0006$

Standard Model \rightarrow



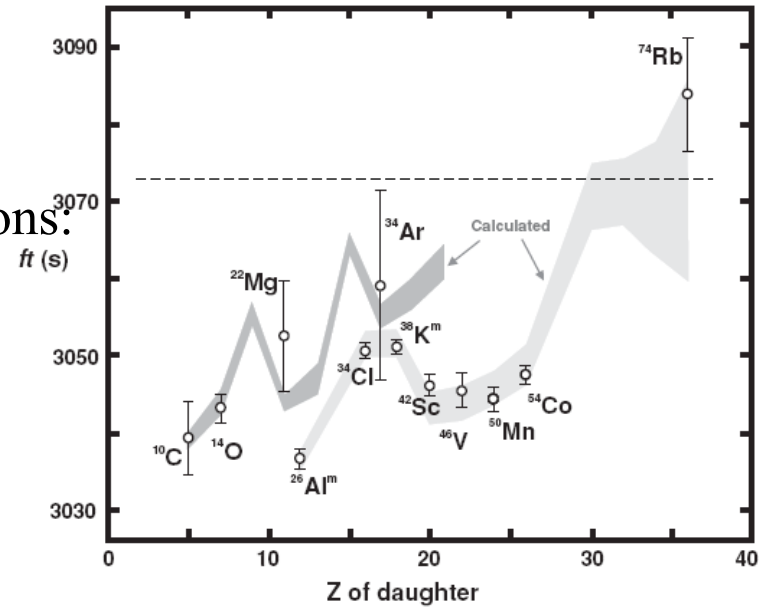
Nuclear super-allowed $0^+ \rightarrow 0^+$ β -transitions

$$ft = \frac{K}{G_F^2 |V_{ud}|^2}$$

(plus corrections)
with half life t ,
phase space factor f

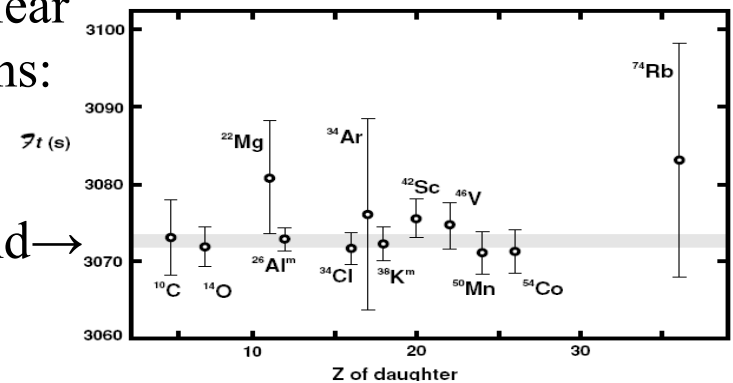
J.C. Hardy, I.S. Towner,
PR C 71, 055501 (2005)

before
nuclear
corrections:



after nuclear
corrections:

1σ band \rightarrow



(from > 100 measurements)

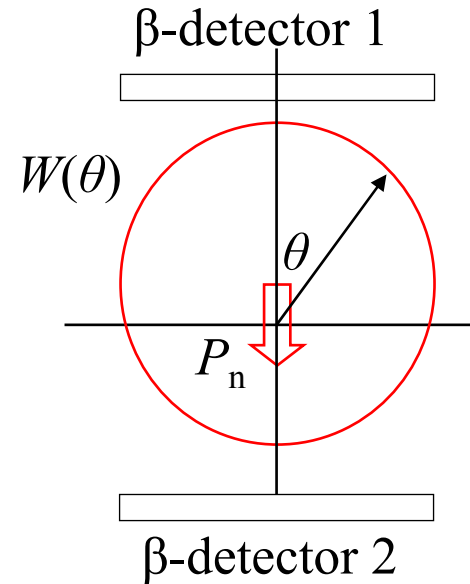
β -asymmetry A in polarized neutron-decay

angular distribution of electrons (Wu-exp't.):

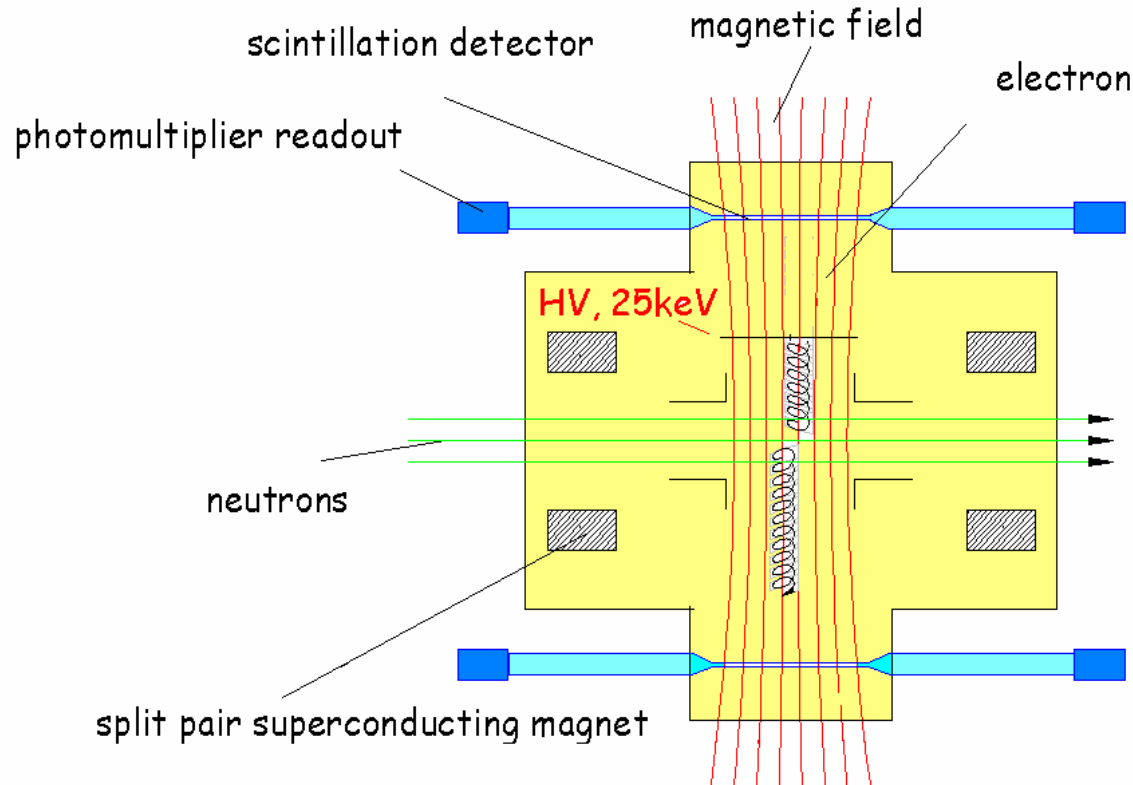
$$W_{\beta}(\theta) = 1 + \frac{v_{\beta}}{c} AP_n \cos \theta$$

polarized cold neutron beam \rightarrow

n-flux $\Phi_n \sim 10^9 \text{ cm}^{-2}\text{s}^{-1}$
n-temperature $T_n \sim 30\text{K}$
n-polarization $P_n = 0.997(1)$

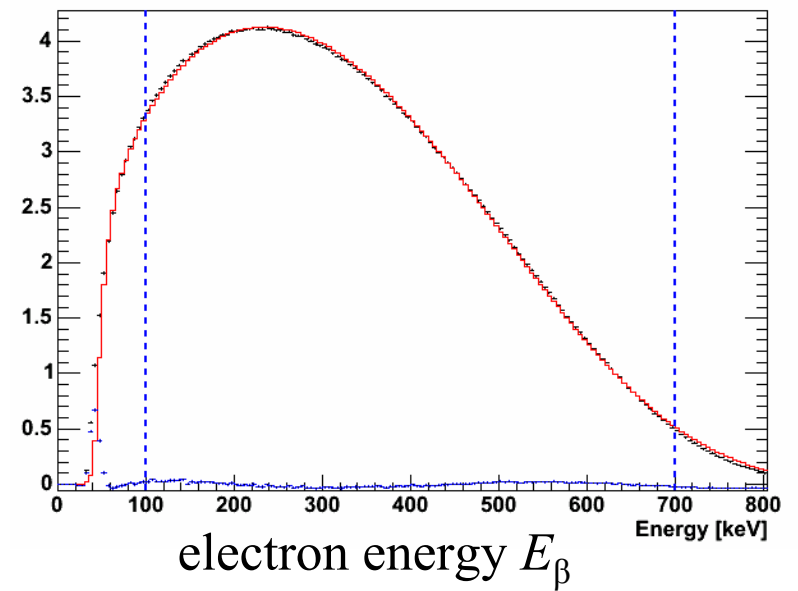
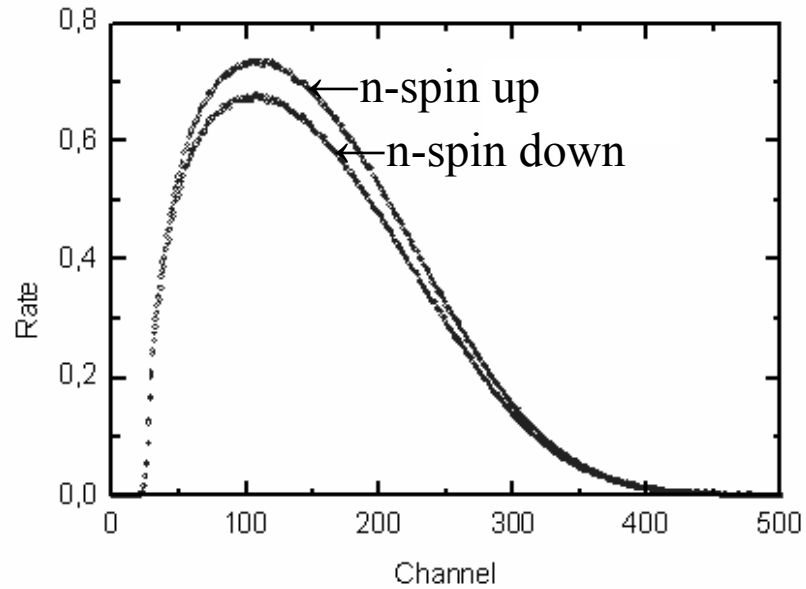


The PERKEO neutron decay spectrometer

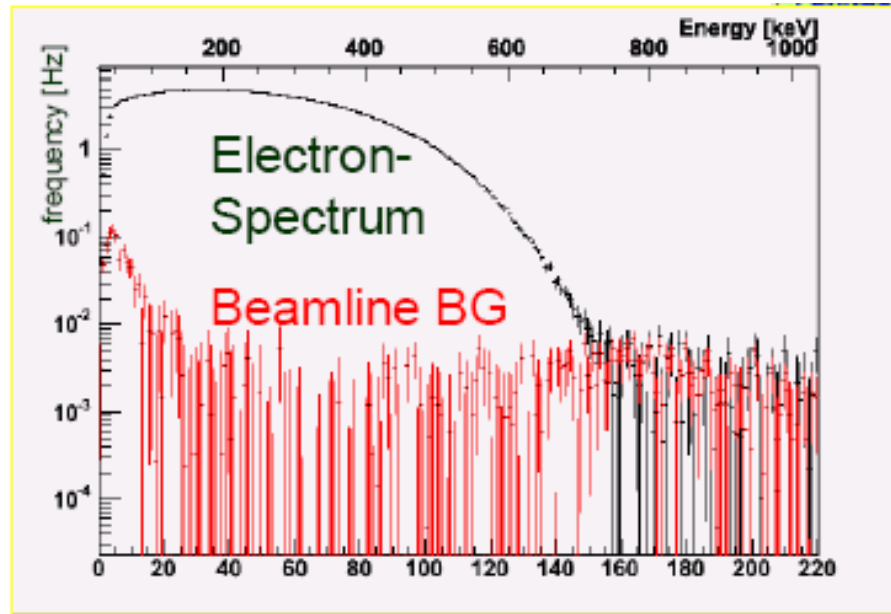


measured: β -asymmetry A
 ν_e '-asymmetry B
proton asymmetry C

β -energy spectra

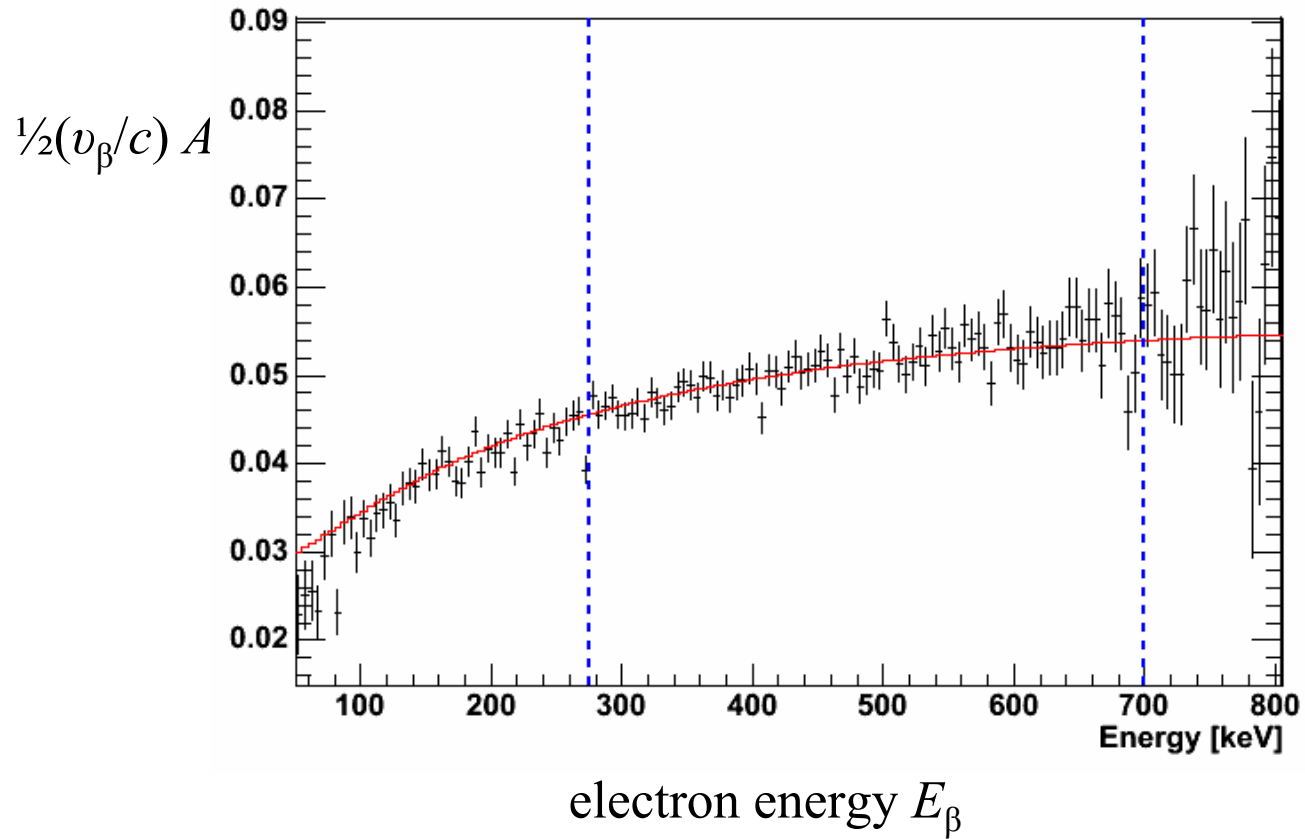


Beam related background



in-beam: 1 of 10^7 neutrons decay in spectrometer;
uncompensated background $< 10^{-3}$

β -asymmetry A :



New Perkeo results (prelim.)

Abele-group

β -asymmetry:

PDG 2006 (≈ 2002): $A = -0.1173(13) \rightarrow \lambda = -1.270(3)$

PERKEO 2002: $A = -0.1189(7) \rightarrow \lambda = -1.274(2)$
 system. corr.: $5 \cdot 10^{-3}$

PERKEO 2006: $A = -0.1195(4) \rightarrow \lambda = -1.275(1)$ **reconfirmed**
 system. corr.: $1 \cdot 10^{-3}$ thesis **Mund 2006** error: statistics domin.

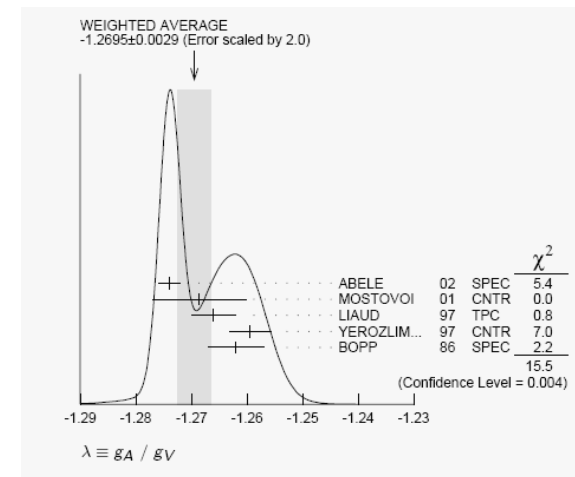
ν -asymmetry:

old world average: $B = 0.981(4)$ dominated by **Serebrov et al. (1998)**

PERKEO: $B = 0.983(5)$ thesis **Schumann 2007**

proton-asymmetry $C = -0.238(11)$ thesis **Kreuz 2004**

$C = -0.239(3)$ thesis **Schumann 2007**



New neutron lifetime

$\tau = (878.5 \pm 0.8)$ Serebrov et al. (2005)

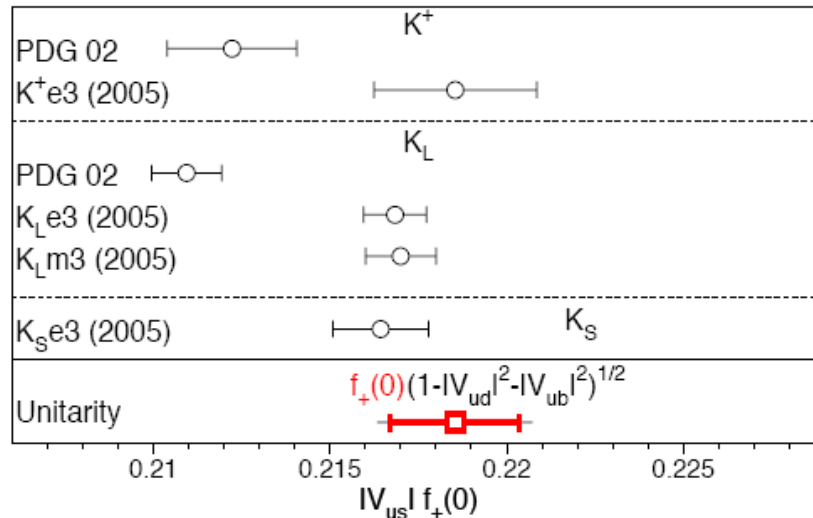
reestablishes unitarity when using old V_{us}

and fits better to cosmology (confirms n_B/n_γ from CMB)

New V_{us} value

= by-product of ϵ'/ϵ -analysis:

2002↓ ↓2005



B.R. $K_L \rightarrow \pi e \nu, \pi \mu \nu$

New CKM element V_{us} reestablishes unitarity when using old τ_n :

PDG 2006, all measurements: $\Delta = 0.0008 (5)_{ud} (9)_{us}$

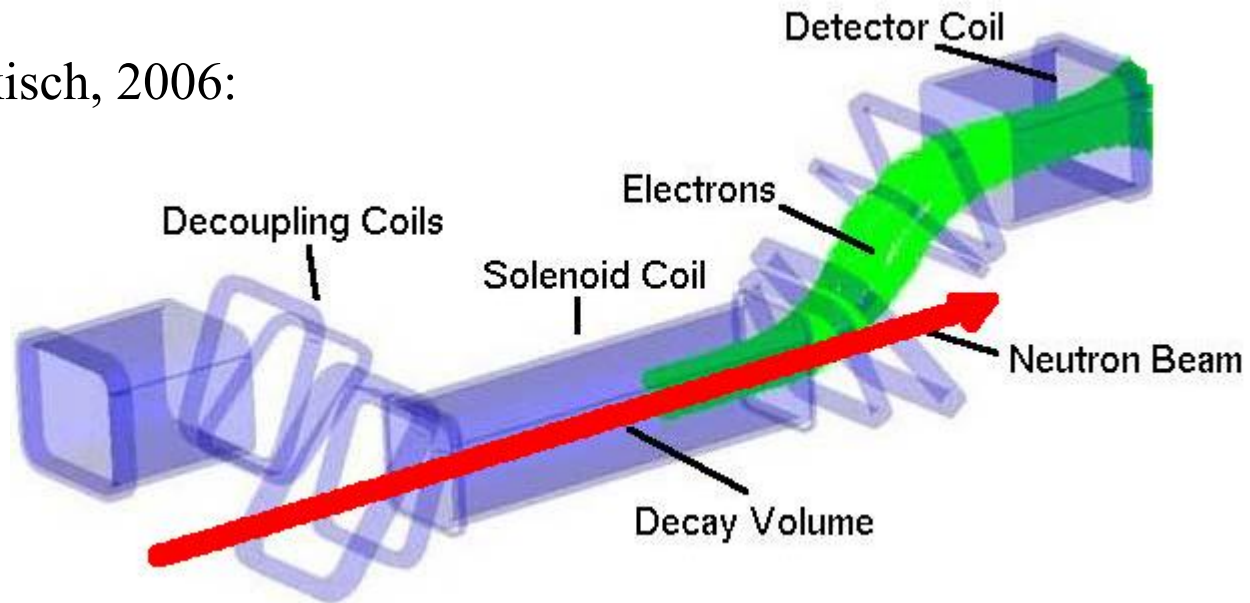
Serebrov, neutron: short τ , high λ : $\Delta = -0.0038 (28)$

- still clarification needed.

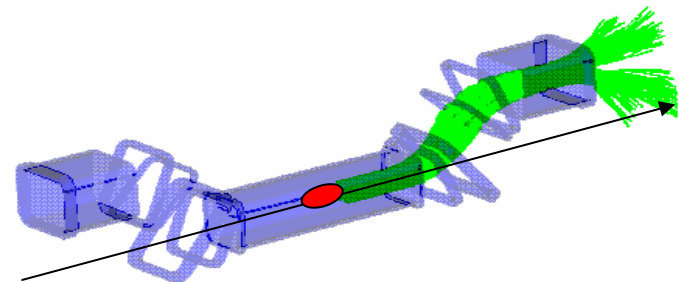
Other strategy: assume unitarity to hold \rightarrow strong-interaction physics

New Perkeo instrument

thesis B. Märkisch, 2006:



later:





PERKEO 3
November 2006
ILL, GRENOBLE

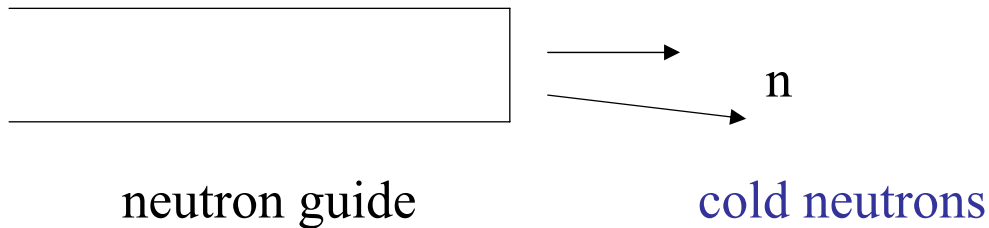
count rate:
25 000 n-decays/sec.
= 100 × PERKEO 2

First measurement:
weak magnetism in n-decay
 $\sim \mu_n - \mu_p$
($\sim 1\%$ effect in β -asymmetry)

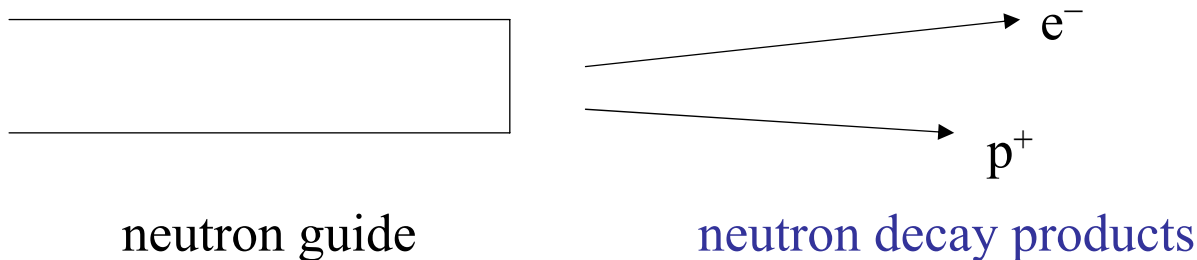
far future:

PERC = strong source of n-decay products ...

conventional neutron beam line:

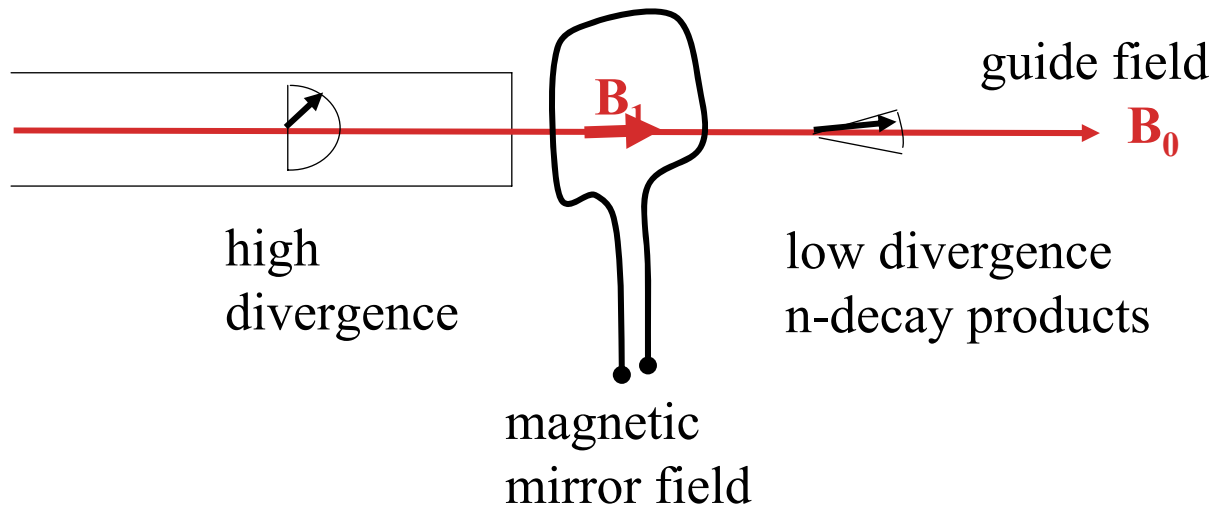


PERC = proton-electron radiation channel:

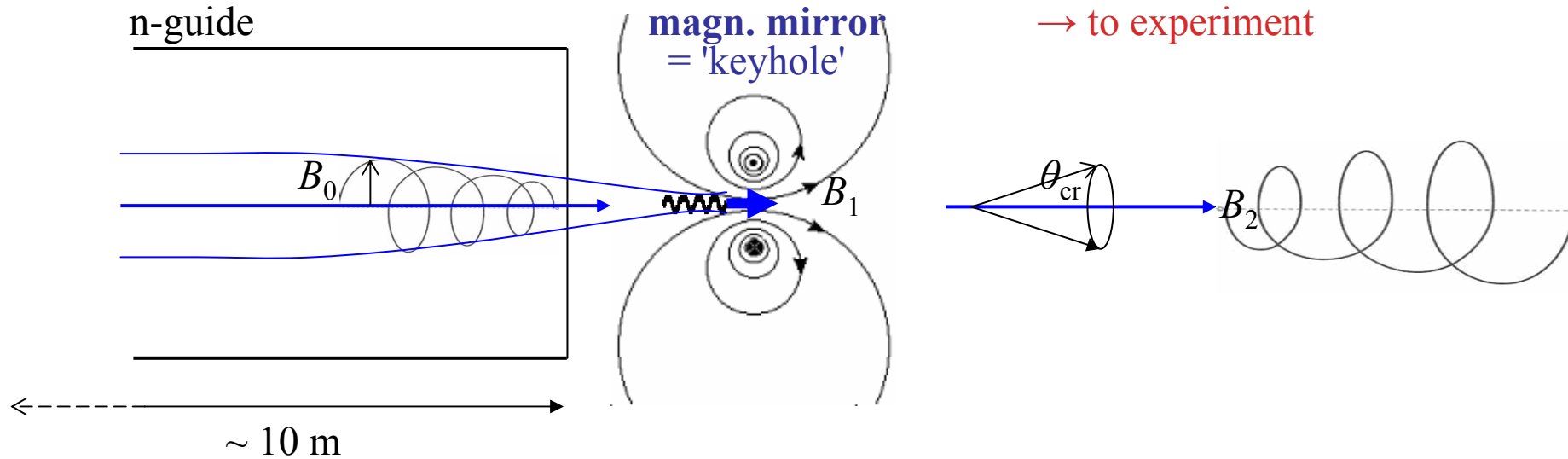


10^6 neutron decays per meter of guide

... of variable beam divergence:



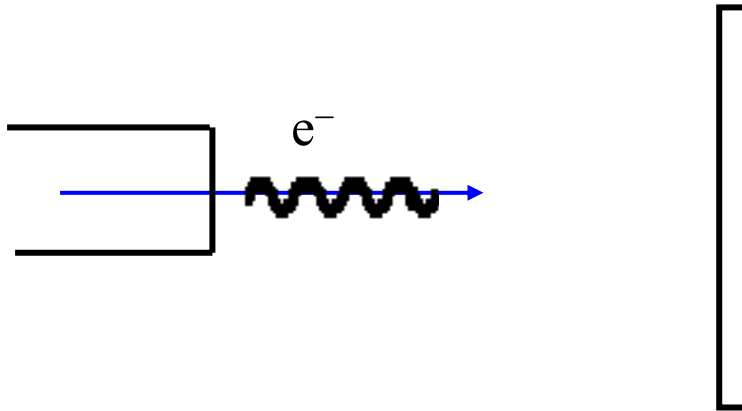
Decay products pass through 'keyhole':



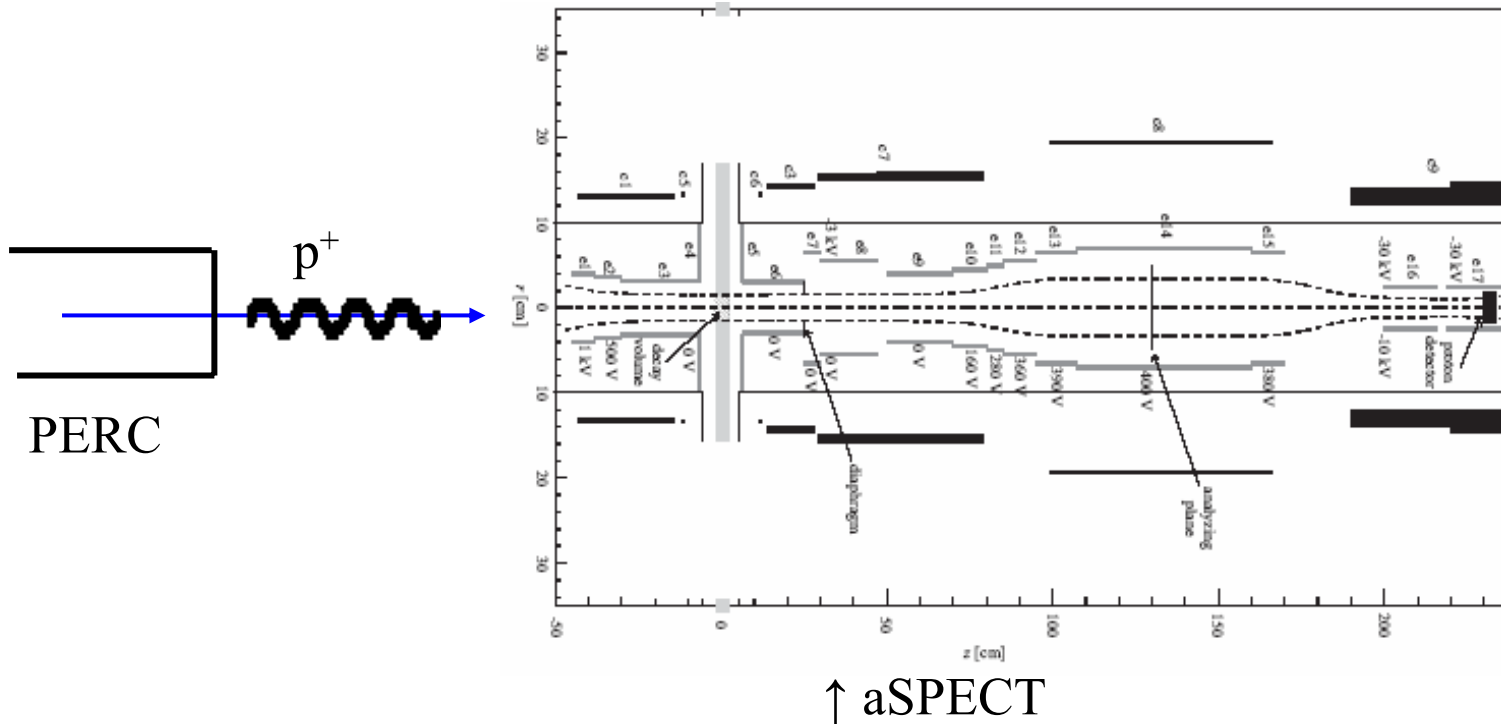
example:

magnetic field:	2 Tesla	8 Tesla	$\frac{1}{2}$ Tesla
gyration radius:	2 mm	$\frac{1}{2}$ mm	4 mm
critical angle:	30°	90°	15°

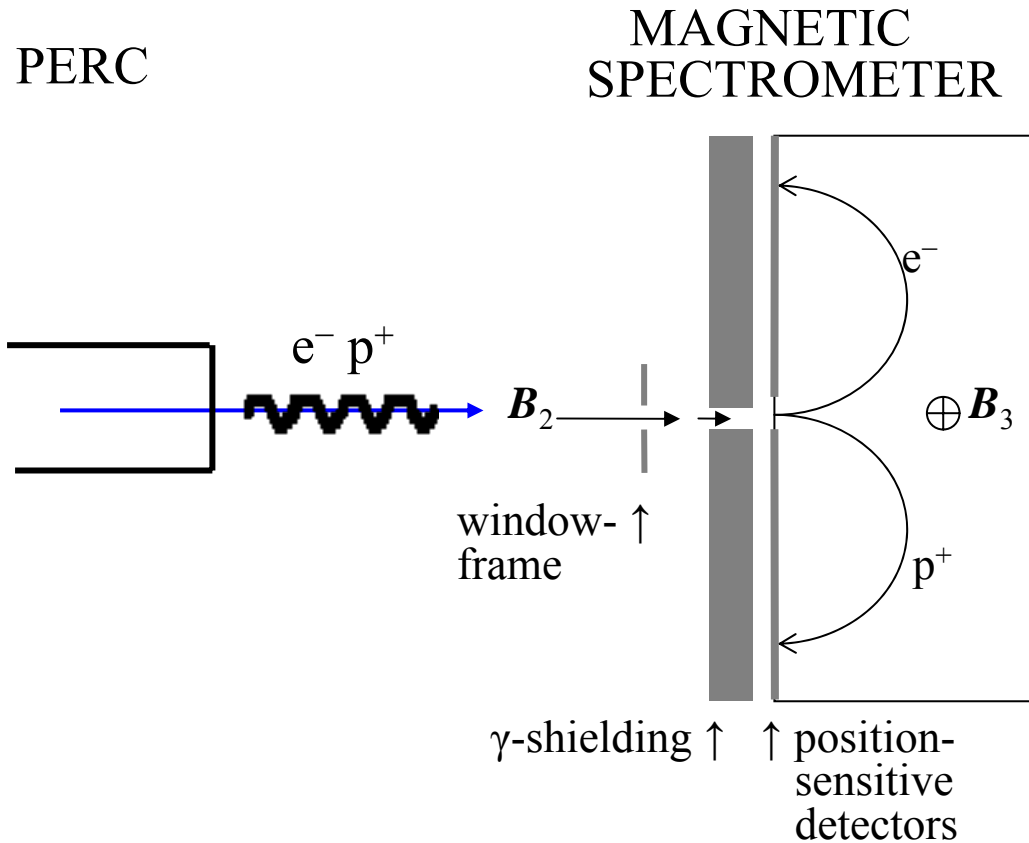
PERC + Perkeo detector



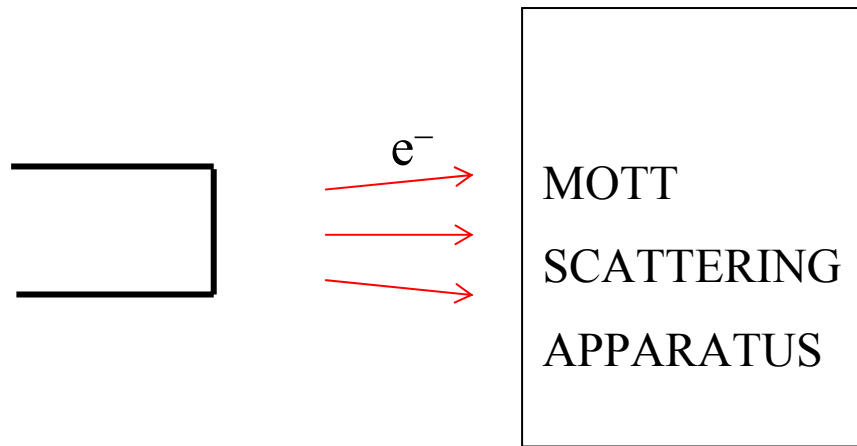
PERC + *a*SPECT retardation spectrometer:



PERC + magnetic p^+ , e^- spectrometer:



PERC + Mott scattering device



test of electron helicity $H_e \sim v_e/c$ in hadron decay

Observables accessible with PERC:

electron and proton spectra N_e, N_p

electron-neutrino correl. coeff. a

beta-asymmetry A

neutrino-asymmetry B

proton-asymmetry C

Fierz interference b

electron helicity H_β

helicity-related correl. coeff.

all in dependence of energy E_β, E_p

or of momentum p_β, p_p or TOF_p

aim: n-decay correlation coefficients to $\sim 10^{-4}$ rel. error
(today: $\sim 10^{-3}$)

2. Neutron oscillations

a) neutron-antineutron oscillations

Neutrinos oscillate:	$\nu_e \leftrightarrow \nu_\mu$, etc.	$\Delta m \cdot c^2$:
Lepton number oscillations	$L_e \leftrightarrow L_\mu$, etc.	= 0.05 eV
Kaons oscillate:	$K \leftrightarrow K'$	
Strangeness oscillations	$S \leftrightarrow -S$	= 10^{-18} eV
Do neutrons oscillate?	$n \leftrightarrow \bar{n}$	
Baryon-number oscillations	$B \leftrightarrow -B$?

Neutron oscillations allowed in various Grand-Unified Theories

The ILL neutron oscillation experiment

The magnetically shielded beam <5nT



The antineutron detector



'Appearance experiment':

Experimental limit

$$\tau_{n \rightarrow \bar{n}} > 0.86 \cdot 10^8 \text{ s (90\% c.l.)}$$

$$m \ll c^2 = \langle n | H | \bar{n} \rangle \approx 10^{-23} \text{ eV}$$

probes 10^5 GeV range (model dependent)

Heidelberg-ILL-Padova-Pavia collaboration (M. Baldo-Ceolin et al., 1994)

b) neutron – mirror-neutron oscillations

is there a sterile mirror world?

Mohapatra, 2005: $n \leftrightarrow n_{\text{mirror}}$

can neutrons spontaneously disappear into sterile,
i.e. unobservable mirror neutrons?

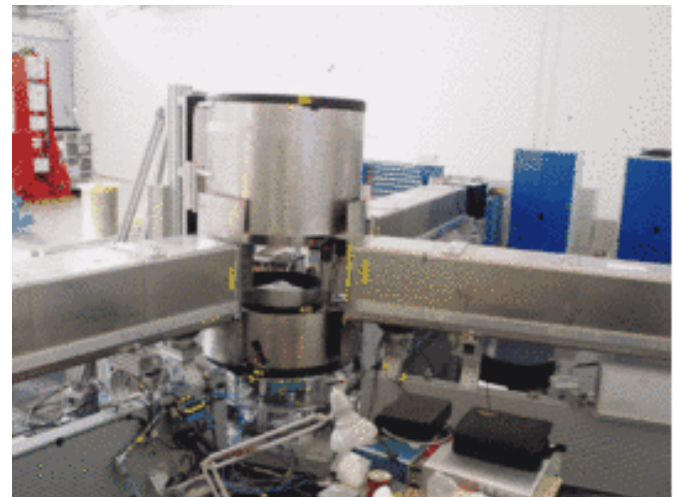
Experiment: U. Schmidt, thesis B. Böhm, 2007:

using zero-field spin-echo apparatus at FRM2, developed by U. Schmidt et al.
and ultrafast 'CASCADE' n-detector, developed by M. Klein, Chr. Schmidt

'disappearance experiment':

experimental limit:

$$\text{rate } (\tau_{n\text{-nmirror}})^{-1} > 2.7 \text{ s}^{-1} \text{ (90\% c.l.)}$$



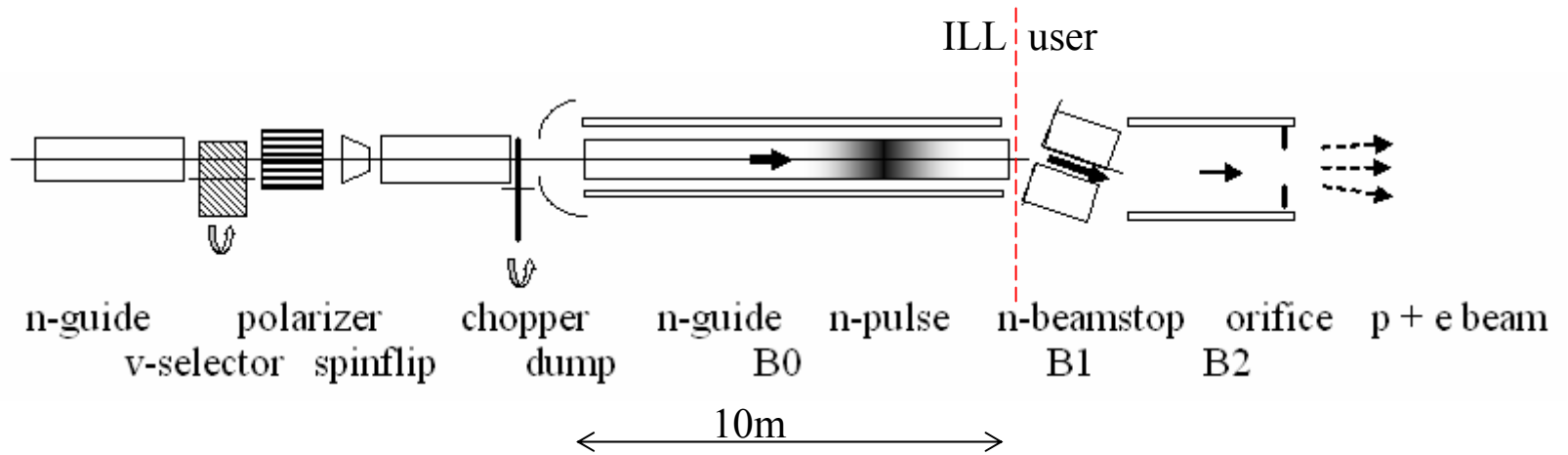
Summary

'Precision Measurements at Low Energy':

Physics very far beyond the Standard Model is best tested:
at low energy in the first family, where precision is highest;
in the electroweak or gravitational sector, where reliable predictions exist.

In neutron decay (as elsewhere) there is still a lot of homework to do.

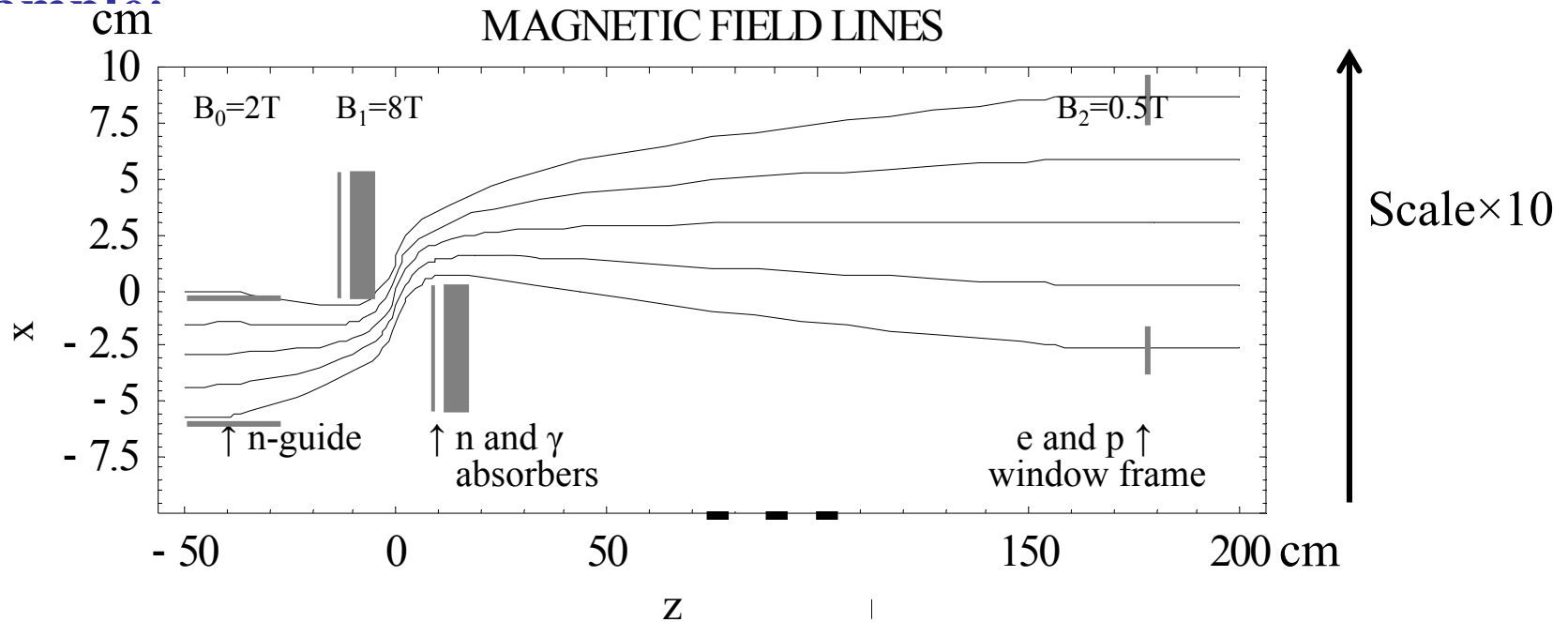
PERC details



PERC neutron beamstop:

Charged neutron decay products can be guided anywhere (electro-)magnetically

Example



Competitors: Neutrons

Europe: by far strongest continuous neutron sources;
 leading in neutron-particle physics
 US: strong new effort
 but: SNS has much weaker continuous neutron beams

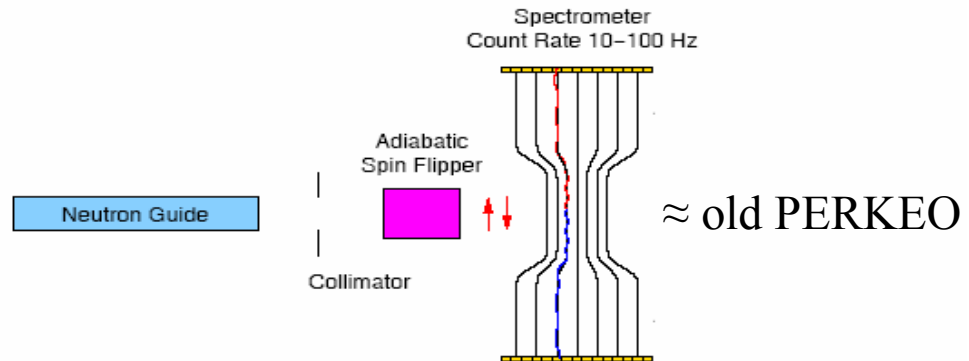
"Fundamental neutron physics-
 instrument development team":

Argonne NL
 Arizona State U.
 BARC
 Cal-Tech
 Depauw U.
 Duke U.
 Harvard U.
 Indiana U.
 Jefferson Lab.
 KEK
 Kentucky U.
 Los Alamos NL
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Michigan U.
 New Hampshire U.
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 N.-Carolina Centr. U.
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 Oak Ridge NL
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