

Particle and Astrophysics: The neutrons perspective

Examples on why to use **neutrons**

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Overview

- The first 3 minutes of the Universe (in a nutshell)
- Weak interaction properties of hadrons
 - Neutron lifetime measurement
- Flavour diagonal CP-violation
 - EDM of the neutron
- Nature and the search for right-handed partners
 - Neutron bound β -decay
- Toolbox: Ultracold neutrons
 - Sources with D_2 converter
 - Sources with superfluid helium
- Summary

History of the Universe

10^{-43} seconds:

The curtain is being lifted...

Space and time are foamlike....

Superstrings: a 'Weltformel' ?

- All forces are unified
- The world is 10-dimensional
- However, only 4 dimensions have participated in the expansion of space All other dimensions are curled up

BIG BANG

10^{-44}
T 10^{-32}
E 10^{19}

Key:

q quark
g gluon
e electron
m muon
n neutrino



History of the Universe

$10^{-35} - 10^{-32}$ s:

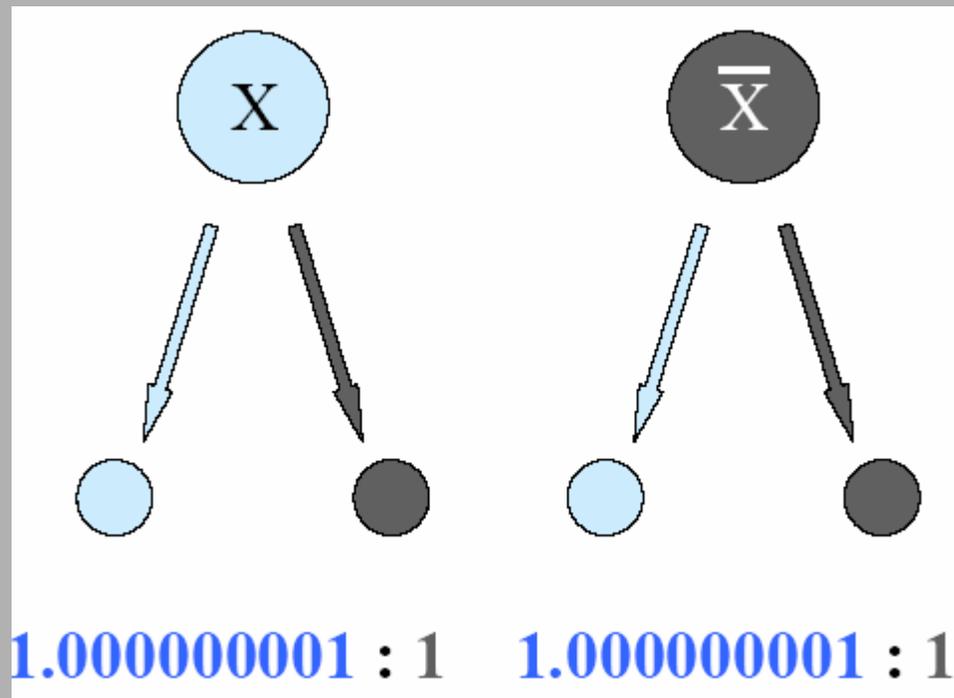
Inflation:

within 10^{-32} seconds space
expands by a factor 10^{50}

- Quantum-fluctuations in the energy density are being amplified
(3 K background radiation)
- The build the base for future galaxy formation

History of the Universe

A tiny excess of matter over antimatter is created



We are made from the 1 in the ninth decimal

History of the Universe

Until 10^{-10} seconds:

Soup from elementary particles and exchange bosons constantly interacting

Thereby cooling off by spatial expansion and modification of coupling strength



History of the Universe

10^{-10} - 10^{-6} seconds: Electroweak symmetry breaking

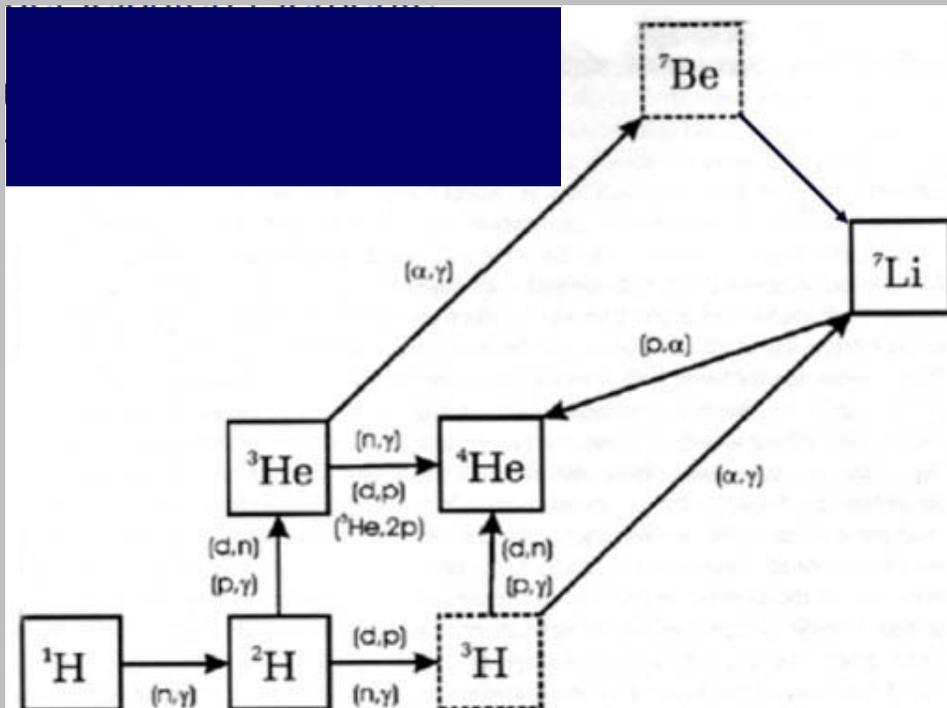
**Neutrinos decouple:
Weak interaction is becoming too
weak, neutrinos stop exchanging
energy and momentum with
other particles**



History of the Universe

10^{-2} - 10^3 seconds:

Creation of first elements:
Primordial nucleosynthesis



No stable elements
with
 $A=5, 8$

${}^4\text{He}$ is heaviest
element

The first 3 minutes
are over



History of the Universe

10^5 years:

Creation of atoms
Decoupling of photons

From 10^9 years on:
Creation of stars and galaxies



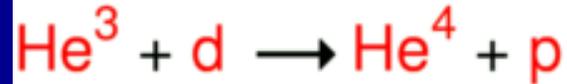
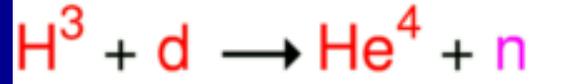
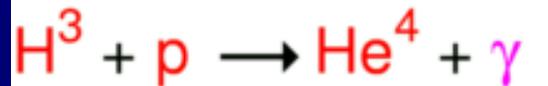
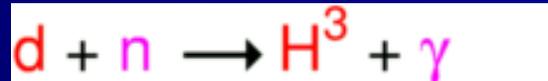
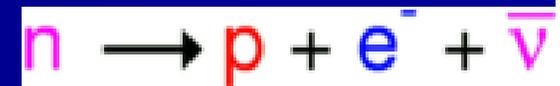
NGC 1850 • Star Clusters in the Large Magellanic Cloud
Hubble Space Telescope • WFPC2

NASA, ESA and M. Romanisho (European Southern Observatory) • STScI PR01-25



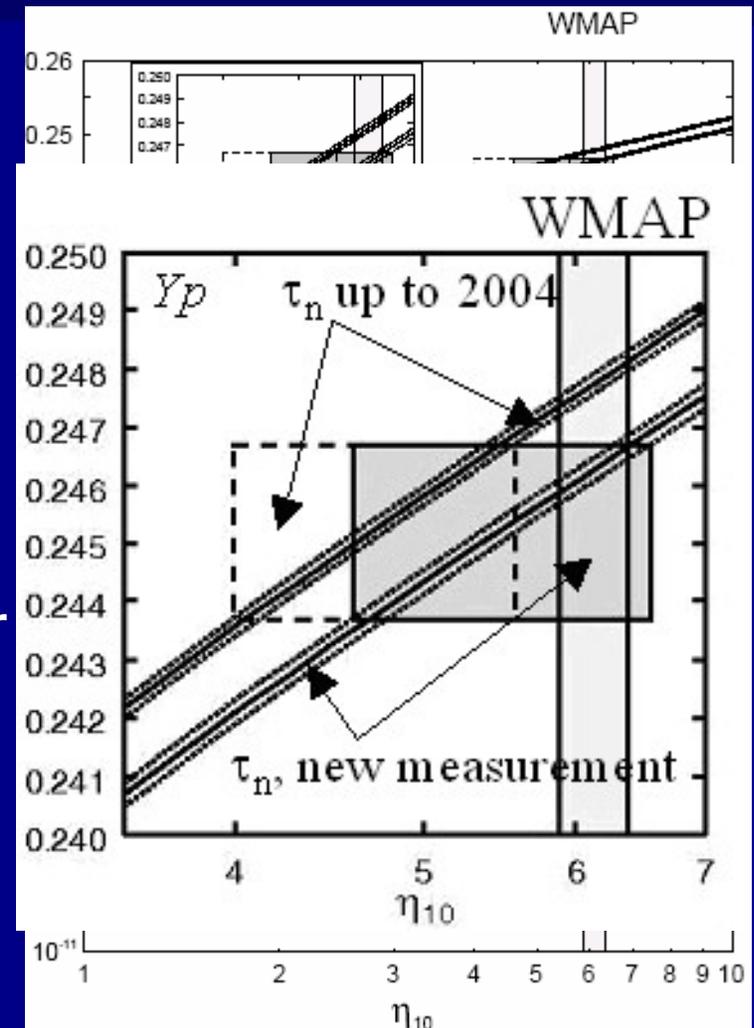
Neutron decay and BB-Nucleosynthesis

- < 1 second after Big Bang:
 - Electroweak process: g_V, g_A
 - proton/neutron = 6:1, freeze out at $kT \sim \Delta m_{pn}$
- > 1 second after Big Bang:
 - proton/neutron changes by **n-decay** to 7:1
- 100 seconds ($kT = 0.1$ MeV)
 - Deuteron formation
- Helium formation

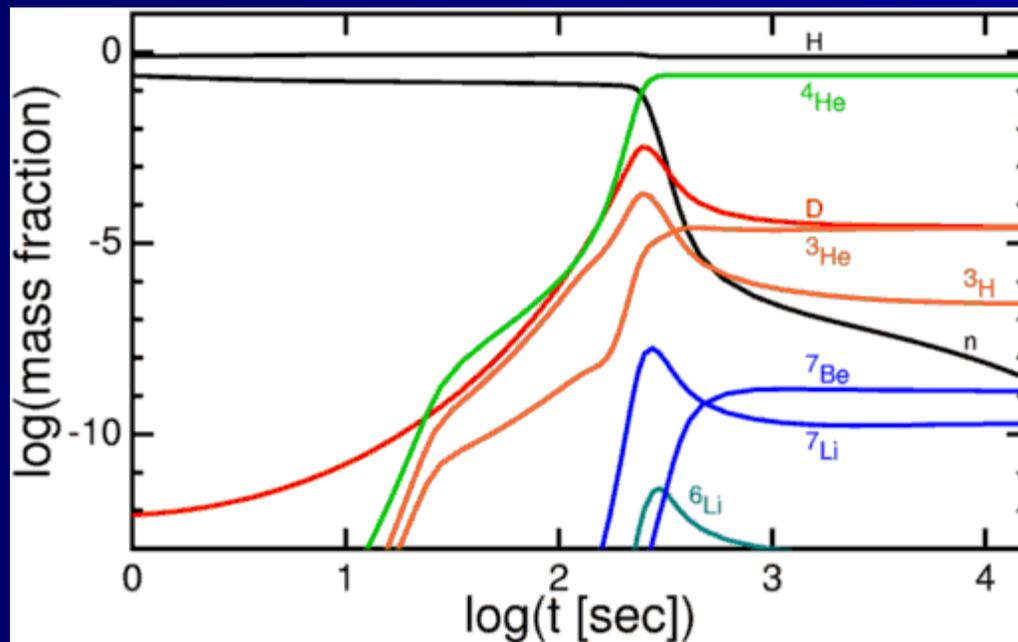


Neutron lifetime and Cosmology

- Three parameters:
 - $\eta_{10} = (n_B / n_\gamma) * 10^{10}$
 - WMAP
 - $Y_p = \text{He} / \text{p}$
 - Metal-poor stars/galaxies
 - τ_n
 - Experiments
- Using standard weak/nuclear physics codes
 - Deuteron abundance (small)
 - Helium abundance
 - Lithium abundance (small)



Primordial Nucleosynthesis



Element generation as function of time

Current status of results

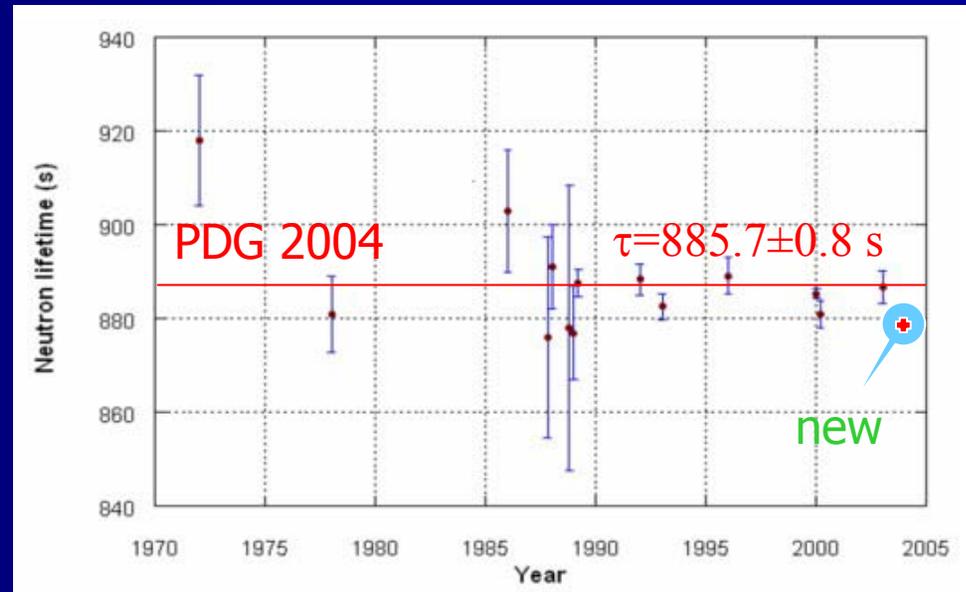
- Lifetime τ : $\tau_n \propto \frac{1}{g_V^2 (1+3\lambda^2)}$ with $\lambda := \frac{g_A}{g_V}$
- Current type of Experiments
 - In beam experiments - count decays in fiducial volume
 - Storage experiments - count decays after a preset storage time t
 - Storage losses
 - Loss of time information

- Present values

$$\lambda = -1.2695 \pm 0.0029$$

$$\tau_n = 885.7 \pm 0.8 \text{ s}$$

[PDG 2006]

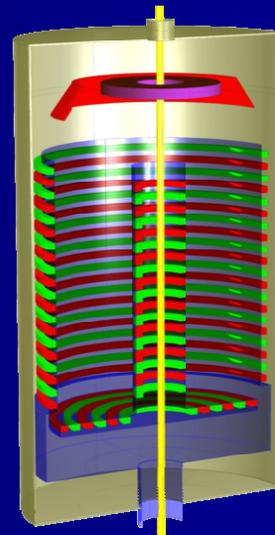


Measurement of the Lifetime

- A **new lifetime** experiment
 - **High statistics** (large volume)
 - Minimize **systematic** effects
 - No wall losses (no material walls)
 - Detection of surviving n after storage time t
 - **Online detection of decay-protons**
 - Detect **defects**
 - Depolarized neutrons
 - Vary spring tension
- **Magnetic bottle** for UCN

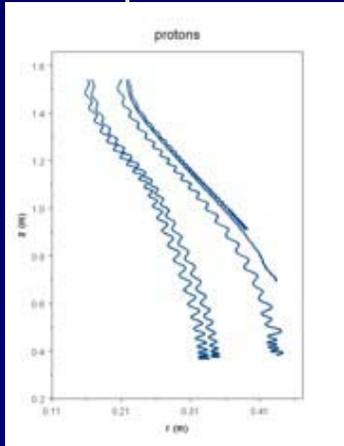
$$F = -\nabla(\boldsymbol{\mu} \cdot \mathbf{B})$$

- $\mu_n = -60.3 \text{ neV/T}$
- $B_{\text{max}} = 2 \text{ T}$
- attraction for $\vec{\mu} \parallel \vec{B}$



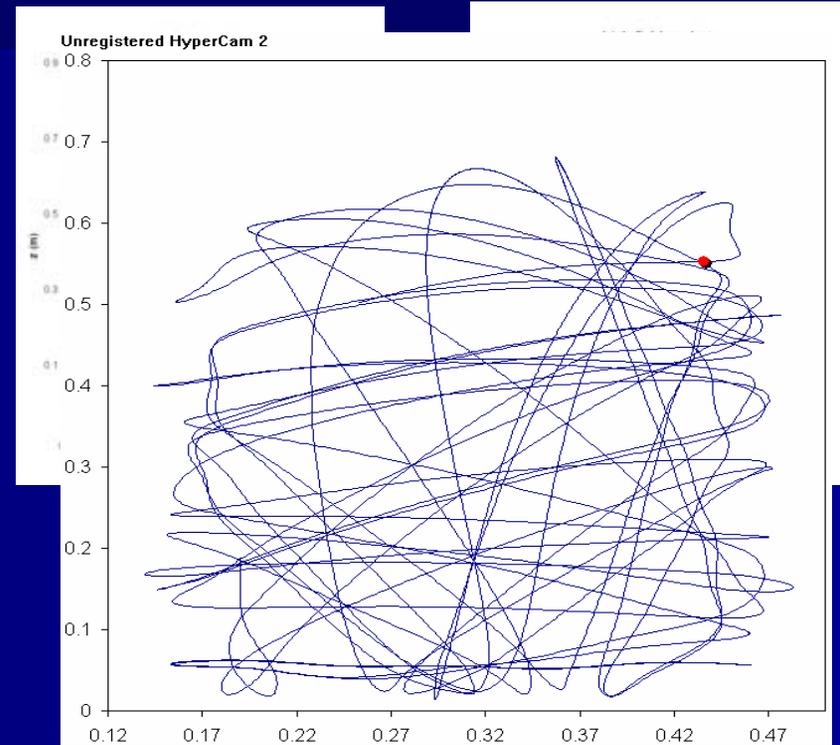
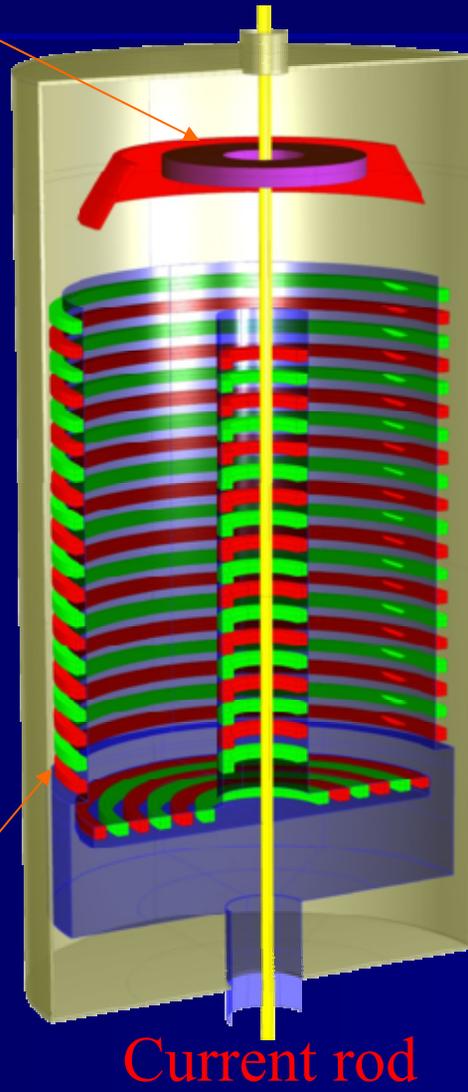
Experimental setup

Proton detection

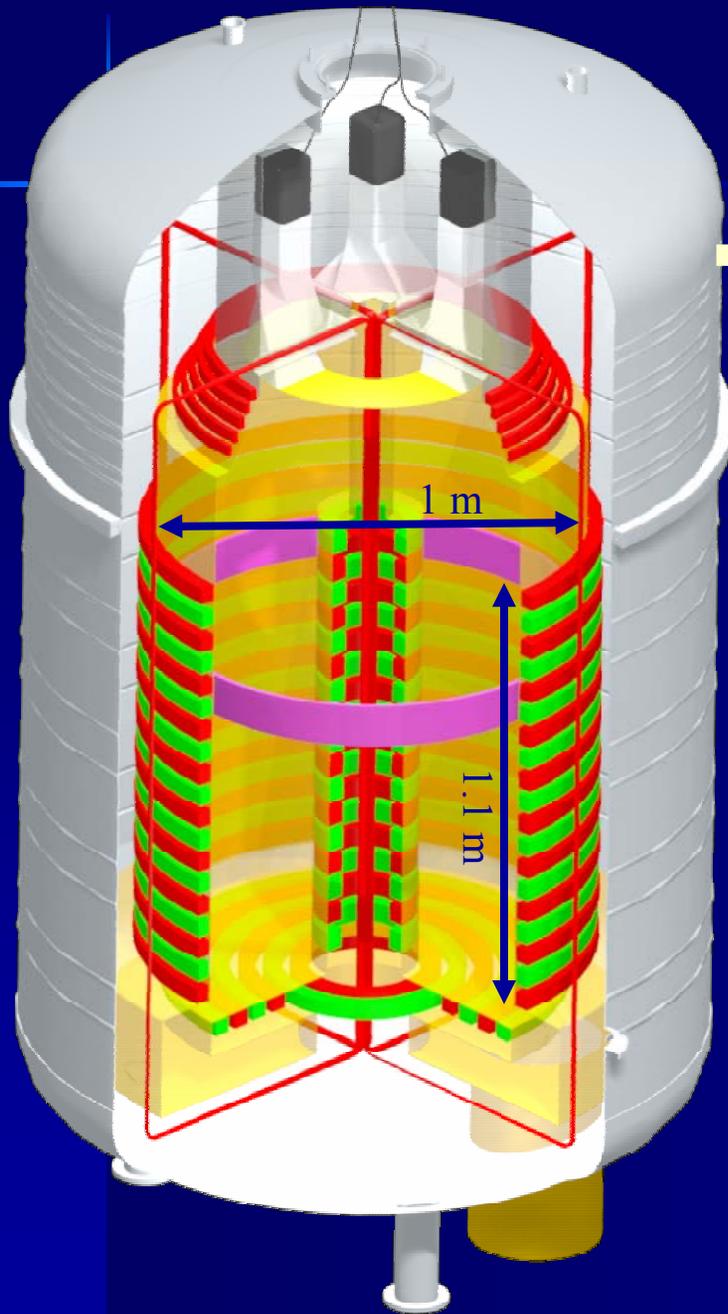


Proton trajectories

Magnetic multipole
 $B_{\text{eff}} \sim 2\text{T}$



Design



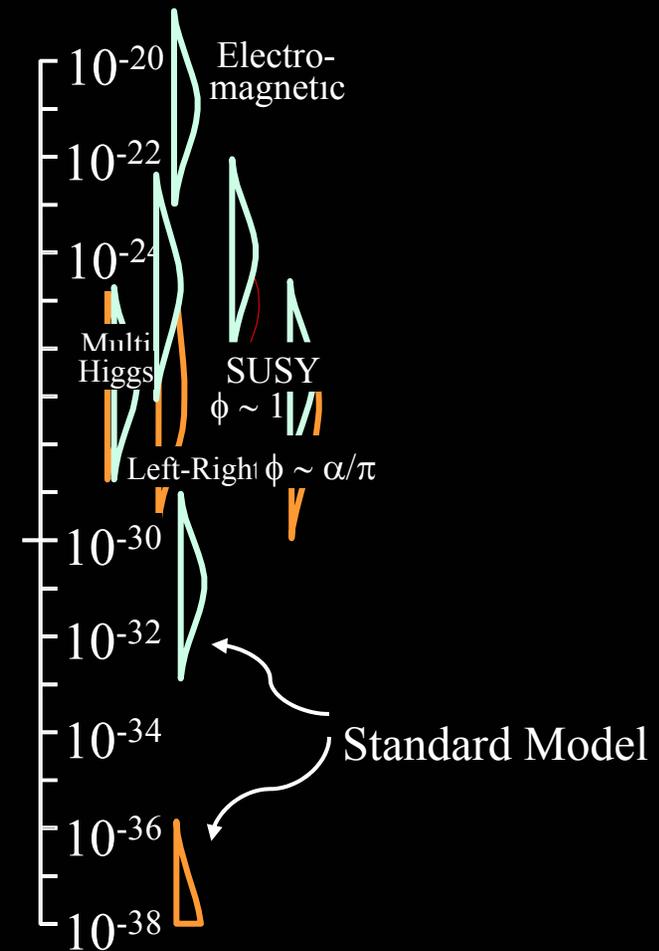
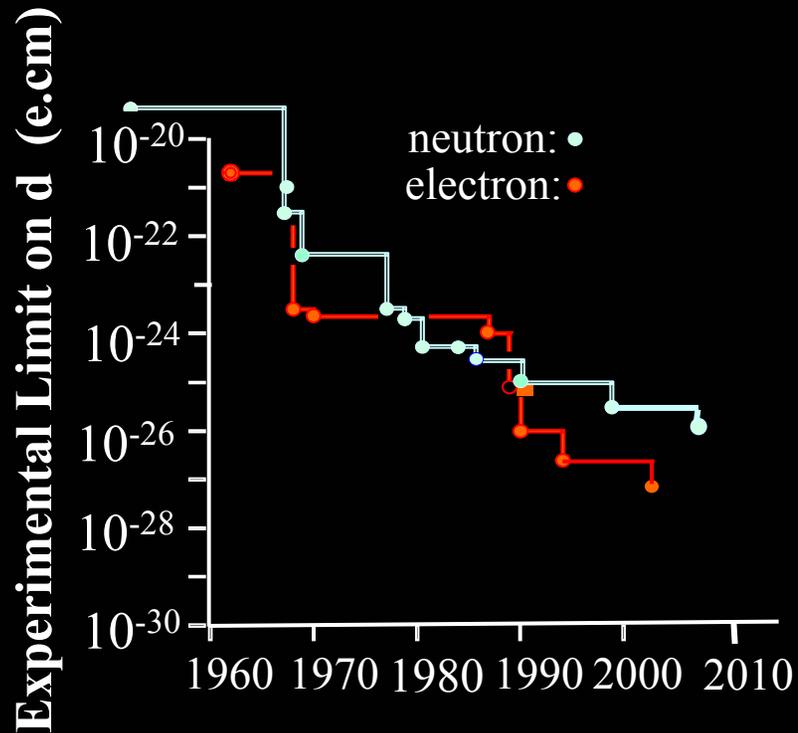
Assume: **new high-density UCN source** (FRMII, PSI)

- UCN (gas-) density: $r = 10^3\text{-}10^4 \text{ cm}^{-3}$
- $B_{\text{max}} = 2 \text{ T}$ $B_{\text{min}} = 10\text{-}3 \text{ T}$
- Volume: 700 l
- $N_{\text{stored}} = 10^7\text{-}10^8$
- **Statistical accuracy:**
 - $\Delta t \sim 1\text{s}$ per measurement cycle (30 min):
 - $\Delta t \sim 0.1\text{s}$ in 2-4 days

CP-violation and EDM

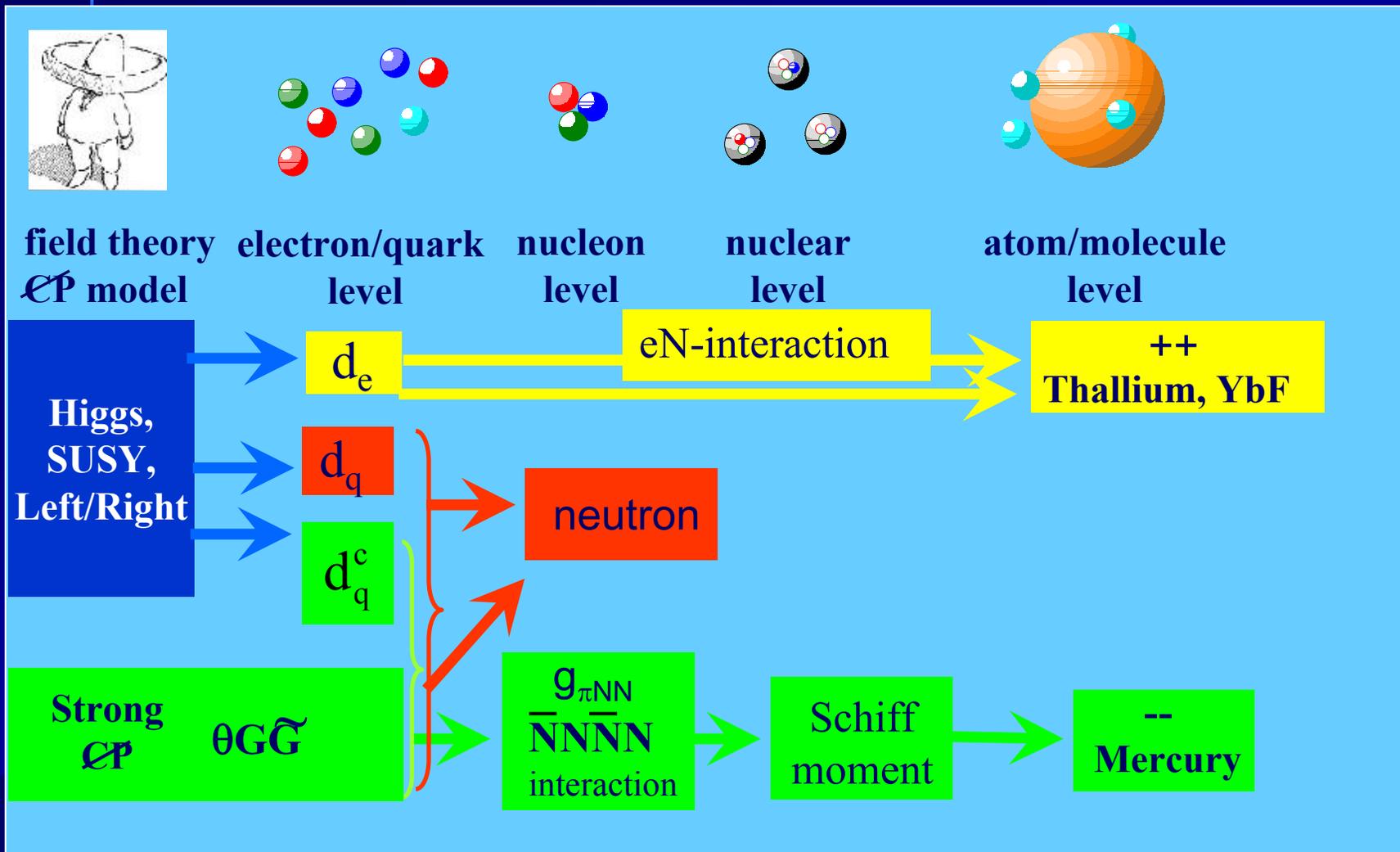
- EDM is a test for flavour diagonal \cancel{CP}
 - Test of structure of vacuum at small distance scales
 - Background free probe for 'new physics' (unlike CKM induced \cancel{CP})
- CP violation in nucleons (neutron) needed for
 - Baryogenesis problem (matter vs. antimatter in the universe)
 - Cosmological requirement (Sakharov criteria)
 - Inflationary scenario suggests dynamical generation of baryon number
 - Test CP violating part in QCD (θ -term)
 - Magically fine-tuned to zero
- EDM can be studied in
 - Diamagnetic atoms
 - Paramagnetic atoms, molecules (\cancel{CP} inducing electron-EDM d_e)
 - Neutron (\cancel{CP} in quark sector)

Search for Electric Dipole Moments



Electric Dipole Moments and Physics

\mathcal{CP} from particles to atoms (main connections)



How to measure an EDM

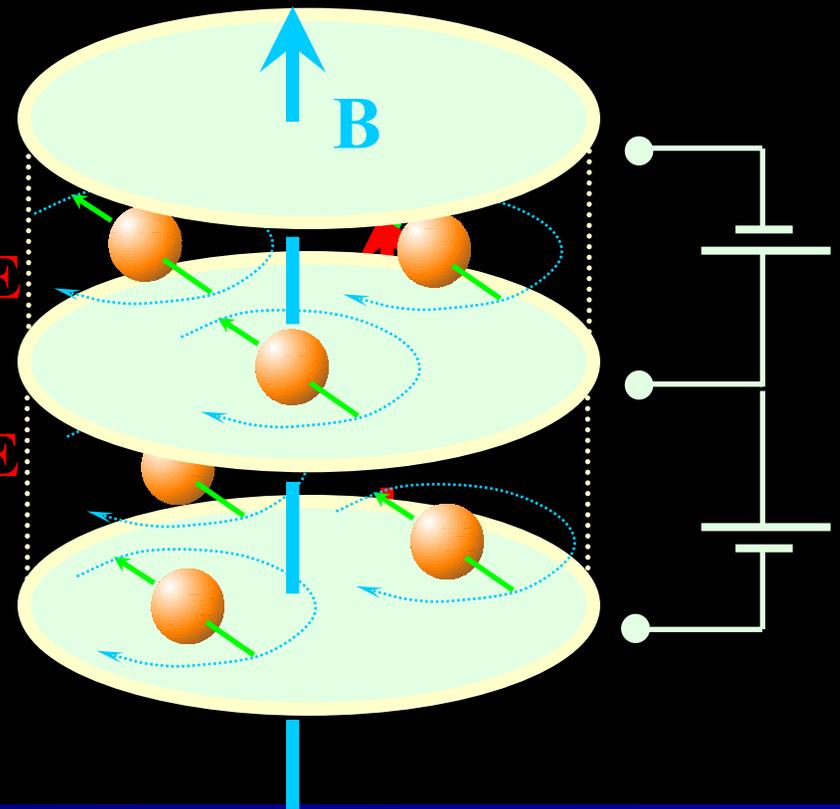
- Ramsey method (split oscillatory fields)
- Two-chamber system

$$\Delta\omega = \omega_{\uparrow\uparrow} - \omega_{\uparrow\downarrow} = 4 \cdot d_n \cdot E/\hbar$$

$$\hbar\omega = \pm \mu B$$

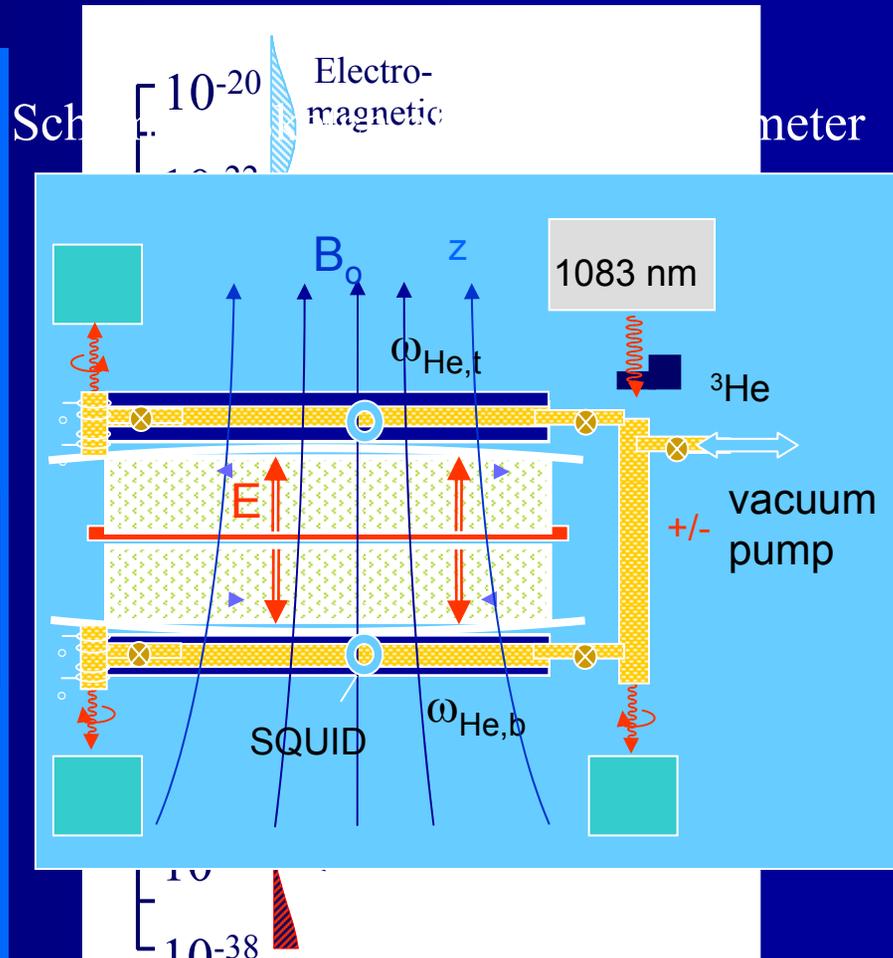
+ dE

- dE



Measurement Accuracy

- Present **limit**: $d_n < 3 \cdot 10^{-26}$ e·cm
- Limitations for δd_n :
 - # stored neutrons
 - New UCN source (e.g. FRMII)
 - E-field strength
 - New material for storage bottle
 - B-field stability (**10fT**)/uniformity (**3pT/cm**)
 - New **magnetometry** (^3He)
 - **shielding**
 - Systematic correlations
 - Field shapes, simulations
- **Aim**: $d_n \sim 10^{-27}$ e·cm (first round limit)
- **Final aim**: $d_n \sim 10^{-28}$ e·cm



$$\sigma_v = 6.5 \cdot 10^{-8} \text{ Hz} \quad \delta B = \sigma_v / \gamma_{\text{He}} \approx 2 \text{ fT}$$

$$\langle \omega_{t,\text{He}} \rangle - \langle \omega_{b,\text{He}} \rangle \propto \left\langle \frac{\partial B}{\partial z} \right\rangle$$

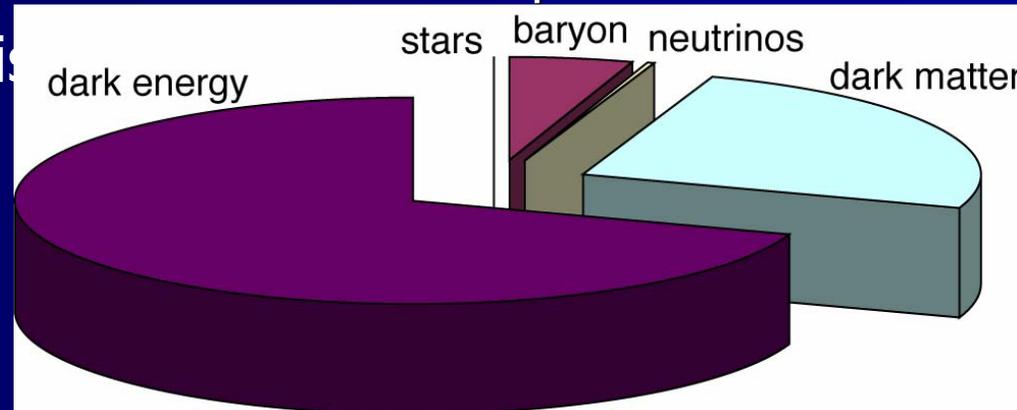
Competition

- Much efforts around the world with different approaches
 - Cryo-EDM at Grenoble (ILL) - super-thermal ^4He source with internal experiment
 - Lamoreaux/Golub in USA (NIST/LANL) - super-thermal ^4He , internal ^3He magnetometer, n-detection... very ambitious and difficult)
 - EDM in Switzerland (PSI) – (multi-cell system)
 - EDM at FRMII – ($\delta B \sim 2\text{ft}$) ! (Mainz – TUM – Gatchina)
- All efforts aim at similar accuracy ($d_n = 10^{-28} \text{ e}\cdot\text{cm}$)
 - Different source techniques
 - Different magnetometry
 - Different systematic effects

Right Handed Neutrinos

- Observation of **Neutrino oscillations** requires neutrino mass
- Popular scenario:
 - **light left-handed** neutrinos
 - **heavy right-handed** neutrinos
 - Seesaw mechanism
- Implications on **cosmology**
 - Neutrino decay as source for Leptogenesis and CP violation
 - Candidate for dark matter particles

- What is the **dark matter** sector ?



The Structure of Weak Interaction

- Standard (V-A) structure of weak interaction embedded in $SU(2)_L$
- Extension

- **Right handed** currents (left-right symmetric models)

- W_R, ν_R

- Measure left-handedness of the ν

- **Tensor** or **scalar** forces

- g_T, g_S

A COMBINED analysis of circular polarization and resonant scattering of γ rays following orbital electron capture measures the helicity of the neutrino.

- Use neutrino

- n

- Small branching fraction

We have carried out such a measurement with Eu^{152, ν_e} which decays by orbital electron capture. If we assume the most plausible spin-parity assignment for this transition with HFS analysis (BRs = 41(0.6%), 13(1.0%), 10(0.2%) (86%, 13, 10% 2s) that the neutrino is "left-handed," i.e., $\sigma_\nu \cdot \hat{p}_\nu = -1$ (negative helicity).

$$\cdot \vec{\sigma}_{\bar{\nu}_e}$$

Measurement Technique

- Unpolarized n decay in magnetic field within reactor tube

- $\vec{p}(H) = -\vec{p}(\bar{\nu}_e)$ with $|E_{\text{kin}}^H| = 326.5 \text{ eV}$

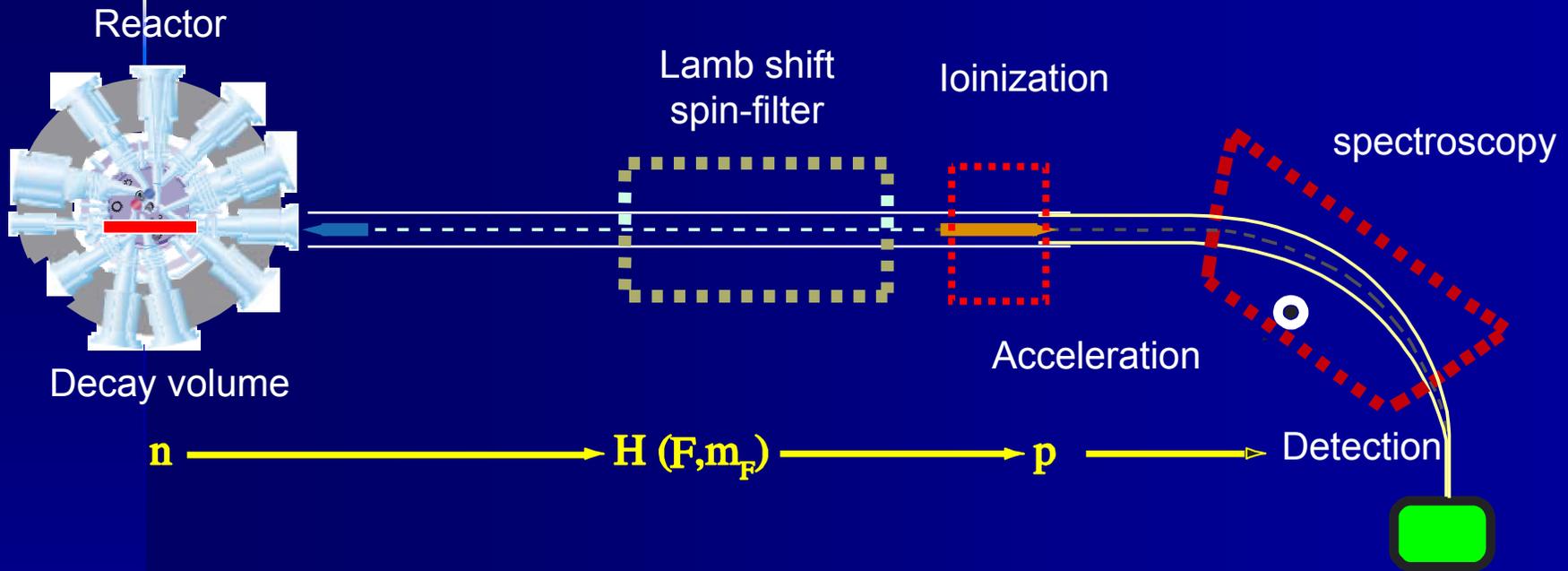
- $\sigma_H = \frac{\vec{\sigma} \cdot \vec{p}}{|\vec{\sigma}| \cdot |\vec{p}|} = 0, \bar{\nu}_e$

- **Select** F, m_F of emerging hydrogen atom using spin-filter method
- **Identify** hydrogen from n-decay via
 - Doppler shifted laser-ionisation process
 - Magnetic spectroscopy
- **Rate**: 0.3 H-atom/s in 2s-state

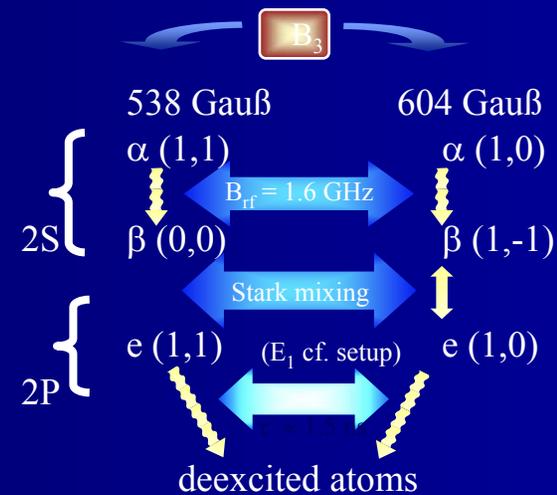
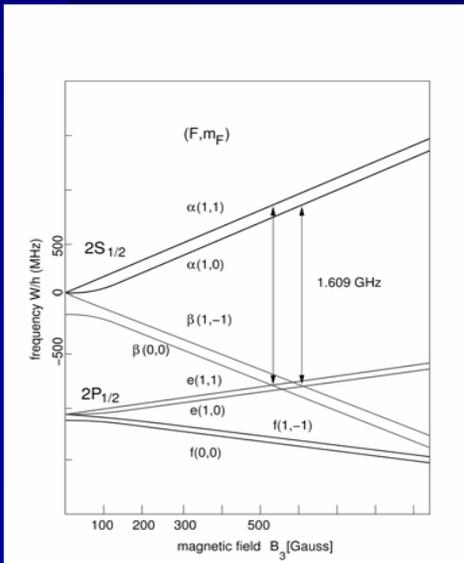
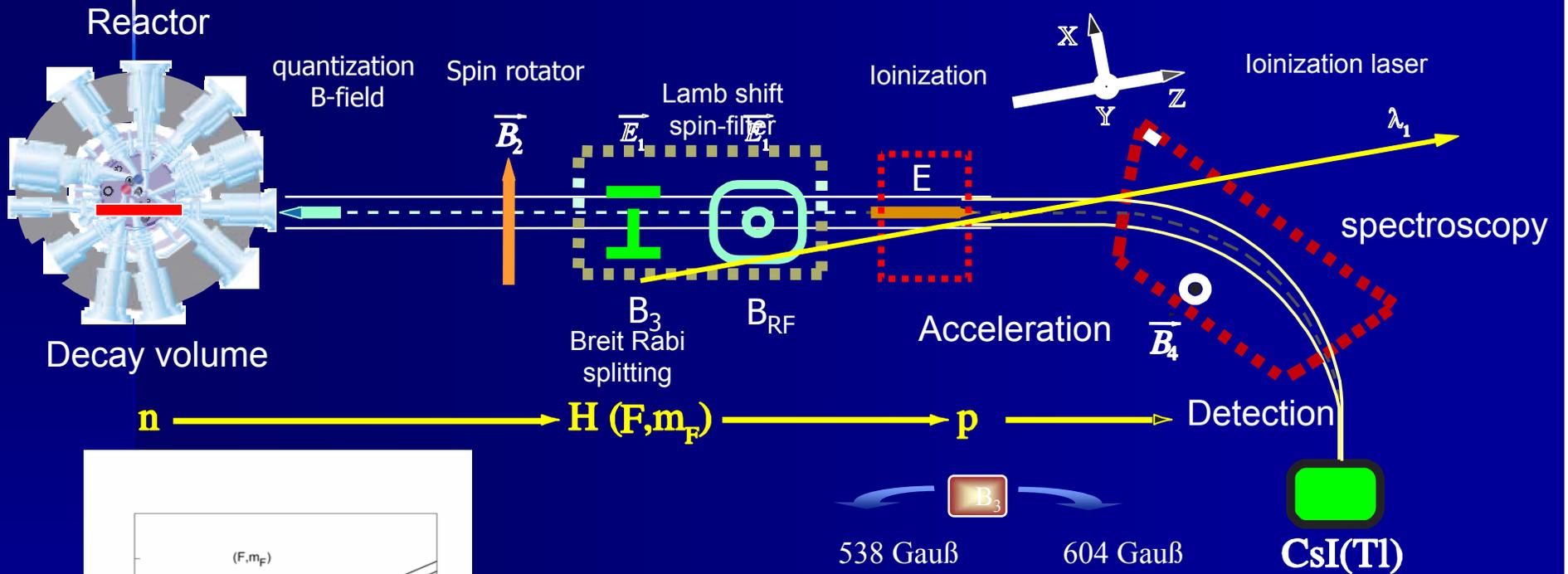
- Physics:

- Relative rates of $F=0,1, m_F=0,1$ give signature for g_S and g_T
- Rate of $F=1, m_F= -1$ shows (V+A)

Experimental setup



Experimental setup



Precision and Competition

- Precision expected:
 - Improvement on g_S upper limit: factor 10 in 2 days/ ε (ε =efficiency)
 - Present accuracy: $|g_S/g_V| < 0.067^*$
 - Improvement on g_T upper limit: factor 20 in 2 days/ ε
 - Present accuracy: $|g_T/g_A| < 0.09^*$
 - Improvement on κ_V : factor 100 in 30 days/ ε (statistically)
 - Present accuracy: 15% from μ, τ decays
- Competition
 - Neutron decay correlations
 - Direct searches for W_R at LHC
 - Muon and tau decay (Michel parameter) – presently best limit

*Severijns et al. 2006: global fit w/o τ_n^{new}

Working Plan n Bound- β -Decay

- Principle: Hydrogen spectroscopy at a reactor

- First stage (2007-2010):

- Laboratory test A

- Laser ionization of hydrogen with
- Magnetic spectroscopy
- Use H-source partially existing

$$|E_{\text{kin}}^{\text{H}}| = 326.5 \text{ eV}$$

- I. **Measurement**: Detect neutron bound β -decay (FRMII)

- Second stage:

- Laboratory test B

- Test spin-filter method
- Need polarized H source (second spin filter ?)

- II. **Measurement** of HFS population

- Setup of high power IR laser (deexcite nS states with $n > 3$)

- III. **Measurement** of HFS population related to _v

Production of Ultra Cold Neutrons

Reactor Neutrons : $E_0 = \sim 2 \text{ MeV}$



Moderator (heavy water) $T = 300\text{K}$
thermal neutrons $E_0 = 25 \text{ meV}$



in or close to reactor core:

Moderator (Deuterium) $T = 20 \text{ K}$
cold $E_0 = 4 \text{ meV}$



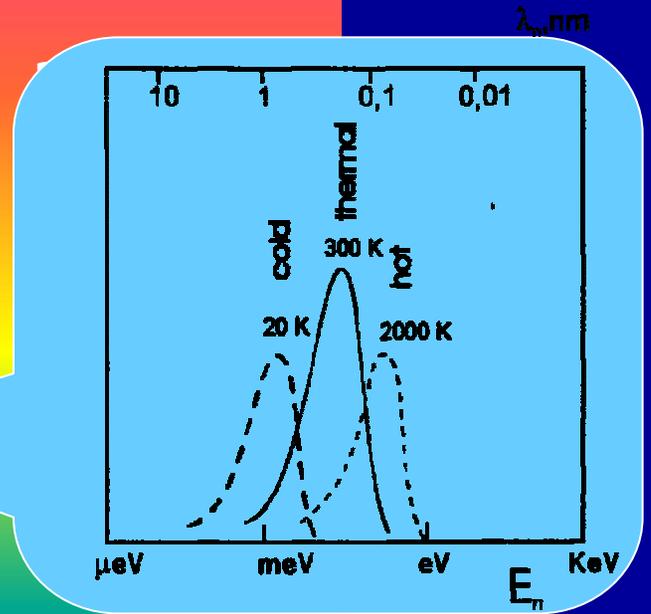
Maxwellian velocity distribution::



Selection and further cooling

Very cold (VCN) $E_0 = 10^{-5} - 10^{-6} \text{ eV}$ 50-15 m/s

ultracold (UCN) $E_0 = 10^{-7} \text{ eV}$ 5 m/s



Superthermal sources

- Two systems are known for production of UCN (non-equilibrium cooling)
 - **Solid deuterium**
 - Upscattering processes < downscattering process
 - At $T \sim 5\text{K}$: $\sigma_{\text{absorption}} \sim \sigma_{\text{upscattering}}$
 - Mean free path of UCN in solid $\text{D}_2 \sim 10\text{-}15\text{cm}$
 - Placement in cold beam w/wo premoderation
 - Position: close to n-source
 - Accumulation mode
 - **Superfluid helium**
 - No upscattering or absorption
 - Usable cold neutron flux for $\lambda \sim 8\text{\AA}$
 - Placement outside strong heat input
 - Typically used for internal experiments
 - Extraction recently demonstrated

Examples of UCN sources

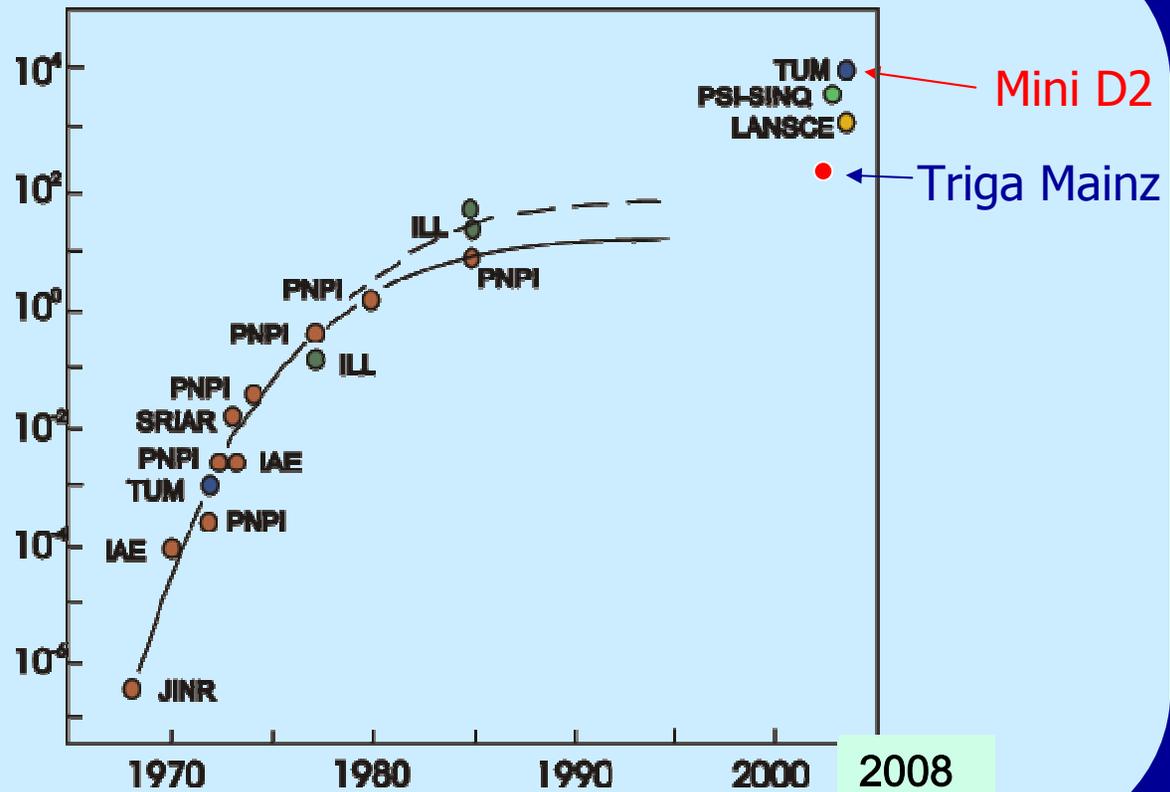
■ Solid D₂- sources:

- LANL
- TRIGA (Mainz)
 - Using TUM geometry/prototype: 100,000 UCN detected in one 'shot'
 - Upgrade in preparation
- PSI (in preparation)
- FRMII (financing now ok)

■ Superfluid helium

- NIST source (built-in lifetime experiment)
- ILL: Cryo-EDM (built-in EDM experiment)
- Small 'portable' test source at FRMII – allows extraction

UCN sources



Summary

- Neutrons offer excellent laboratory to study **particle physics and cosmology**
 - CP-violation and **matter-antimatter asymmetry**
 - Nucleon weak interaction and **primordial nucleosynthesis**
 - Right-handed part of nature – **dark matter** candidates
- Some key experiments performed with **stored neutrons**
 - Technological development of UCN sources
 - Large financial investment based on existing n-sources
 - Technological challenges
- **Precision experiments** required
 - **Systematic and statistical accuracy** mandatory
 - New ideas coming up – exploration stage